Pygmy E1 and giant M1 resonances in the nucleosyntheis of heavy elements

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Outline

- **1.** Basics of the statistical model with emphasis on γ -ray strength function in (n,γ) and (γ,n) reactions
- (γ.n) data indicative of extra γ-ray strengths (pygmy E1 & giant M1)
- 3. HF model predictions of (n,γ) cross sections with γ -ray strength function
- 4. Summary

Collaborators

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Nucleosynthesis of elements heavier than Fe

Leading reactions

Ζ

s, r-processes: neutron capture

p-process: photodisintegration, [proton capture]



Experimentalist's view of the statistical model

Assumption: compound nuclear reactions under ΔE D ΔE : energy spread of incident particles D: average level spacing or high $\rho(U)$

Hauser-Feshbach model A(a,b)B

$$\sigma_{\beta\alpha} = \frac{\pi}{k_a^2} \sum_{J,\pi} g_J \frac{T_{\beta}T_{\alpha}}{T_{tot}}$$

 $\alpha = a + A, \beta = b + B$

statistical factor

$$g_J = \frac{2J+1}{(2J_a+1)(2J_A+1)}$$

Neutron Capture: ${}^{A}X(n,\gamma)^{A+1}X$



 γ -ray strength function $\dot{f}_{E1}(\varepsilon_{\gamma}) = \varepsilon_{\gamma}^{-3} \langle \Gamma_{E1} \rangle / D$

γ-ray transmission coeff. $T_{E1}(\varepsilon_{\gamma}) = 2\pi \langle \Gamma_{E1} \rangle / D$ $= 2\pi \varepsilon_{\gamma}^{3} f_{E1}(\varepsilon_{\gamma})$

$$P_{E1}(E,J,\pi) = \sum_{\nu} T_{E1}^{\nu} + \int T_{E1}(\varepsilon_{\gamma})\rho(E-\varepsilon_{\gamma})d\varepsilon_{\gamma}$$



Enhancements of neutron capture via pygmy E1 and giant M1 resonances



Goriely, PLB (1998)

GDR vs GDR+PDR

CN vs CN+DC



Closed-shell nuclei with $2 \le S_n$ [MeV] ≤ 4

 $1 \le S_n [MeV] \le 3$

DTBA (Discrete Time Blocking Approximation)

Avdeyenkov, Goriely, Kamerdzhiev, Tertychny (2008)

Key ingredients Single particle continuum Phonon coupling



Correct prediction of PDR in ¹³²Sn at 9.8 MeV



Phonon coupling increases (n,γ) cross sections by a factor of 2-3.

AIST Electron Accelerator Facility; National Institute of Advanced Industrial Science and Technology (AIST)



Tsukuba Electron Ring for Acceleration and Storage (TERAS) at AIST



• Energy
$$E_{\gamma} = 1 - 40 \text{ MeV}$$



Neutron Detector System

Triple-ring neutron detector 20^{3} He counters (4 x 8 x 8) embedded in polyethylene



Ingredients in the Talys code

Talys code: Koning, Hilaire, Duijvestijn, Proc. Int. Conf. on Nuclear Data for Science and Technology, AIP Conf. Proc. 769, 1154 (2005).

GDR γ-ray strength function
Lorentzian models: Axel, PR126 (1962), Kopecky & Uhl, PRC41 (1990)
HFB+QRPA model: Goriely, Khan, Samyn, NPA739 (2006)

#Spin-flip giant M1 γ-ray strength functionGlobal systematics in RIPL Handbook (Bohr & Mottelson)Lorentzian function : $E_0=41A^{-1/3}$ MeV, $\Gamma_0=4$ MeV, $f_{M1}=1.58 \ 10^{-9} \ A^{0.47}$ MeV-3 at 7 MeV

Pygmy E1: No global systematics

HFB+ Combinatorial model: Hilaire & Goriely, NPA779 (2006)

M1 strength in Zr isotopes in the photoneutron channel



γ -ray strength function for ^{92}Zr



 96 Zr(γ ,n) 95 Zr

 γ -ray strength functions E1 : HFB+QRPA plus <u>M1 resonance in Lorentz shape</u> E_o = 8.5 MeV (9.0 MeV for ^{91,92,94}Zr) σ_0 = 7.5mb Γ = 2.5 MeV

NLD HFB+ Combinatorial Goriley & Hilaire (2008)

Optical potential Koning & Delaroche (2003)





in 0.01 – 1 MeV

in 0.01 – 1 MeV

Source of uncertainties

NLD models

1.HFB+Combinatorial 2.BSFG **3.CT (Constant Temp.)** 4.GSM (Gen. Superfluid) **5.HFBCS+statisticales**

Optical potential models 1.KD (Koning & Delaroche 2003) 2.JLM (Bauge et al. 2001)

Pygmy dipole resonance in ¹¹⁷Sn

GDR (HFB+QRPA) + PDR $\Box E_0 = 8.5 \text{ MeV}, \Gamma = 2 \text{ MeV}, s_0 = 7 \text{ mb in Gaussian shape} \Box$





Kovaert et al., PRC57 (1998)

γ -ray strength function for ¹¹⁷Sn



Pygmy dipole resonance in ¹¹⁶Sn







Summary

- Here γ strength function is a key nuclear ingredient in the Hauser-Feshbach model calculations of neutron capture rates.
- H is important to investigate the low-energy E1 γ-ray strength function of GDR as a leading factor and pygmy E1 and giant M1 resonances as extra strengths.
- **K** In particular, it is necessary to investigate experimentally and theoretically the nuclear structure and the global systematics of PDR by combining different probes of γ , *p*, ³He, e, and others.

₭ After all, we still have lots of work to do for E1, M1, E2, ...