

Measurement of Strength Function and Level Density of ^{138}La at **OCL** Oslo Cyclotron Laboratory

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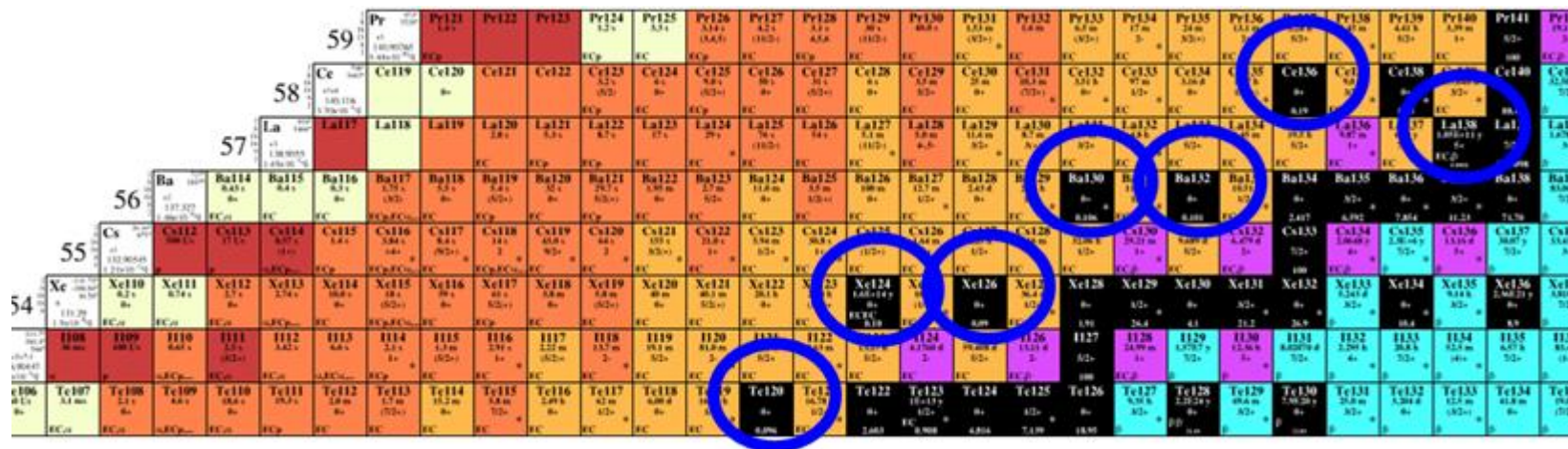
Outline

- » Introduction
- » Experimental Setup
- » Preliminary results
- » Final Remarks

Introduction

The p-nuclei

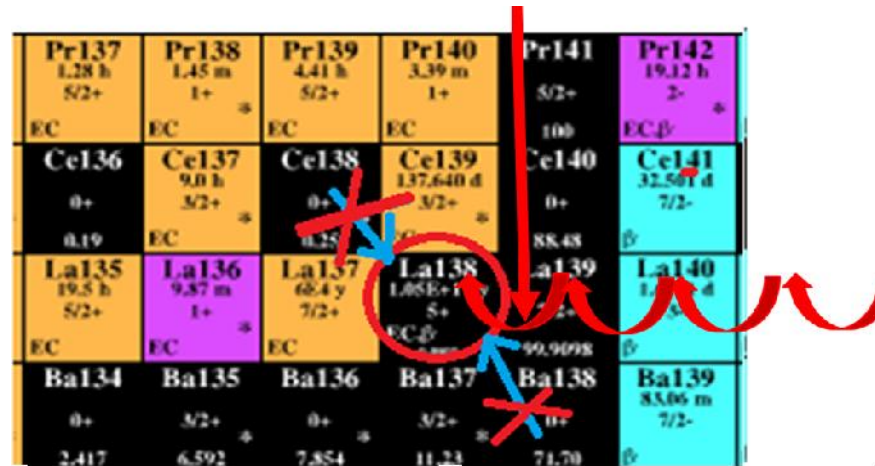
- Most nuclei heavier than ^{56}Fe are synthesized in stars by s- and r- processes
- However, 35 p-nuclei, including ^{138}La , are shielded from s- and r- processes by the valley of stability



- Hence there exist different mechanisms responsible for synthesis of these nuclei
 - The (p,) reaction and / or photo-disintegration
- Very high temperatures and proton density required in this scenario
- Due to high coulomb barrier of heavy nuclei, (p,) is negligible for nuclei with $Z \geq 54$
- By nature rp terminates at Tellurium due to β^+ -decay
- Photo-disintegration remains the main source for the heavier p-nuclei
 - producing very proton rich nuclei via (, n), (,) and (, p) reactions

Introduction

^{138}La is a special case



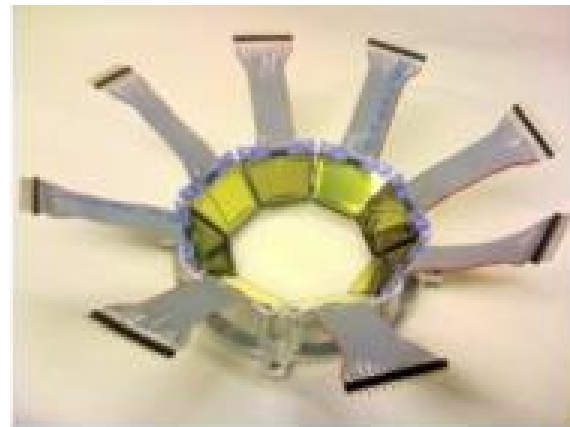
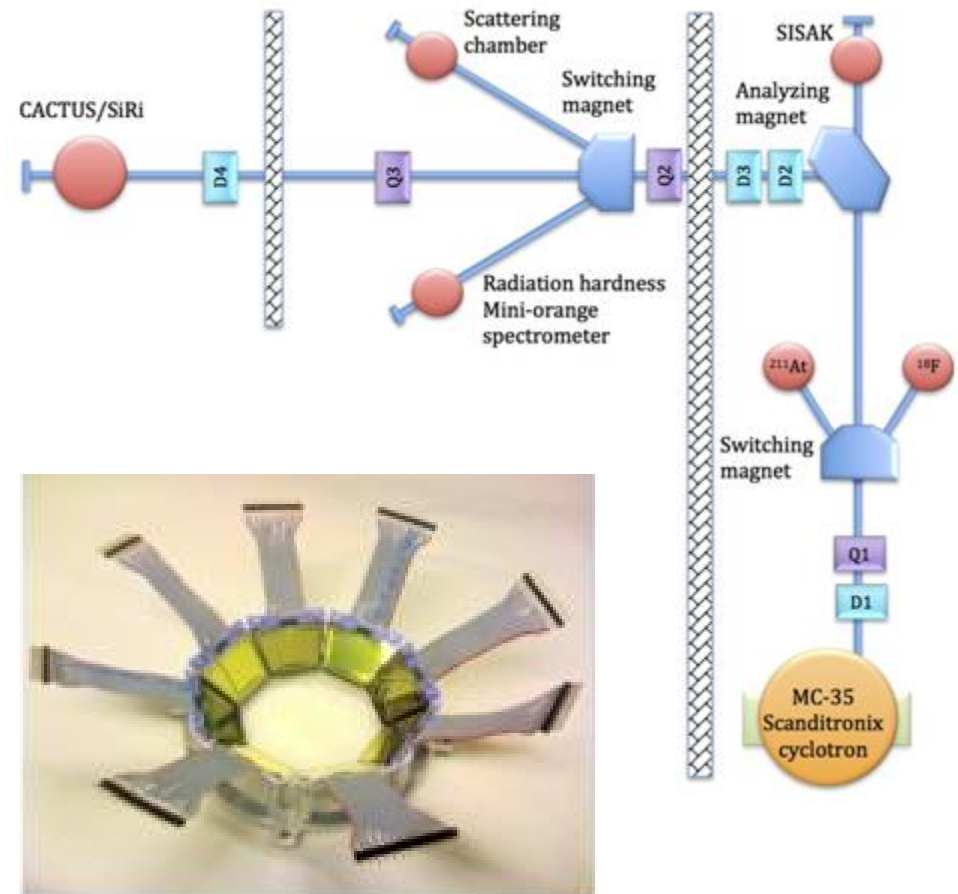
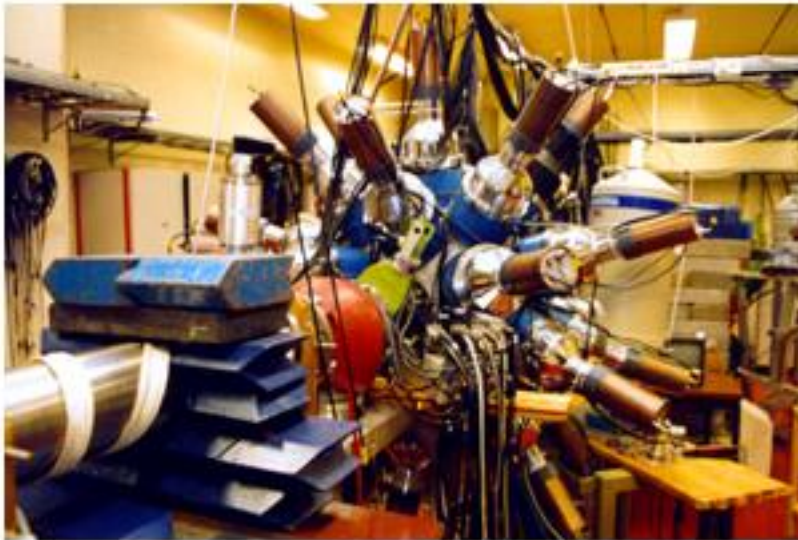
Pr137 1.28 h 5/2+ EC	Pr138 1.45 m 1+ EC	Pr139 4.41 h 5/2+ EC	Pr140 3.39 m 1+ EC	Pr141 5/2+ 100	Pr142 19.12 h 2- ECβ
Ce136 0+ 0.19	Ce137 9.0 h 3/2+ EC	Ce138 0+ 0.25	Ce139 137.640 d 3/2+ ECβ	Ce140 0+ 88.48	Ce141 32.561 d 7/2- β
La135 19.5 h 5/2+ EC	La136 9.37 m 1+ EC	La137 68.4 y 7/2+ EC	La138 1.05E+11 y 5+ ECβ	La139 1.1E+11 y 2+ β	La140 1.1E+11 y 1+ β
Ba134 0+ 2.417	Ba135 3/2+ 6.592	Ba136 0+ 7.854	Ba137 3/2+ 11.23	Ba138 0+ 71.70	Ba139 83.06 m 7/2- β

- Photo-disintegration cannot satisfactorily explain the observed abundance of ^{138}La
- More exotic reactions e.g. $^{139}\text{La}(\nu, n)^{138}\text{La}$ or $^{138}\text{Ba} + \nu e \rightarrow ^{138}\text{La}$ must be considered
- Neutrino induced reactions can explain the observed abundance of ^{138}La
- However (, n) processes cannot be ruled out due to high errors in their rate predictions, which are due to limited knowledge and uncertainties in nuclear level densities, -ray strength function and neutron optical potentials which are critical input parameters for rates calculation

Hence we measured the nuclear level densities and -ray strength function using the **Oslo Method**

Experimental Setup

- Beam time: first half of February 2013
- ^3He beam at 38 MeV
- $^{139}\text{La}(^3\text{He}, ^4\text{He})^{138}\text{La}$



- CACTUS array, 26 NaI γ -ray detectors
- SiRi array, 64 Si telescope detectors (particle identification)
- At $40 - 54^\circ$ with respect to the beam

Experimental Setup

La target and experimental challenges

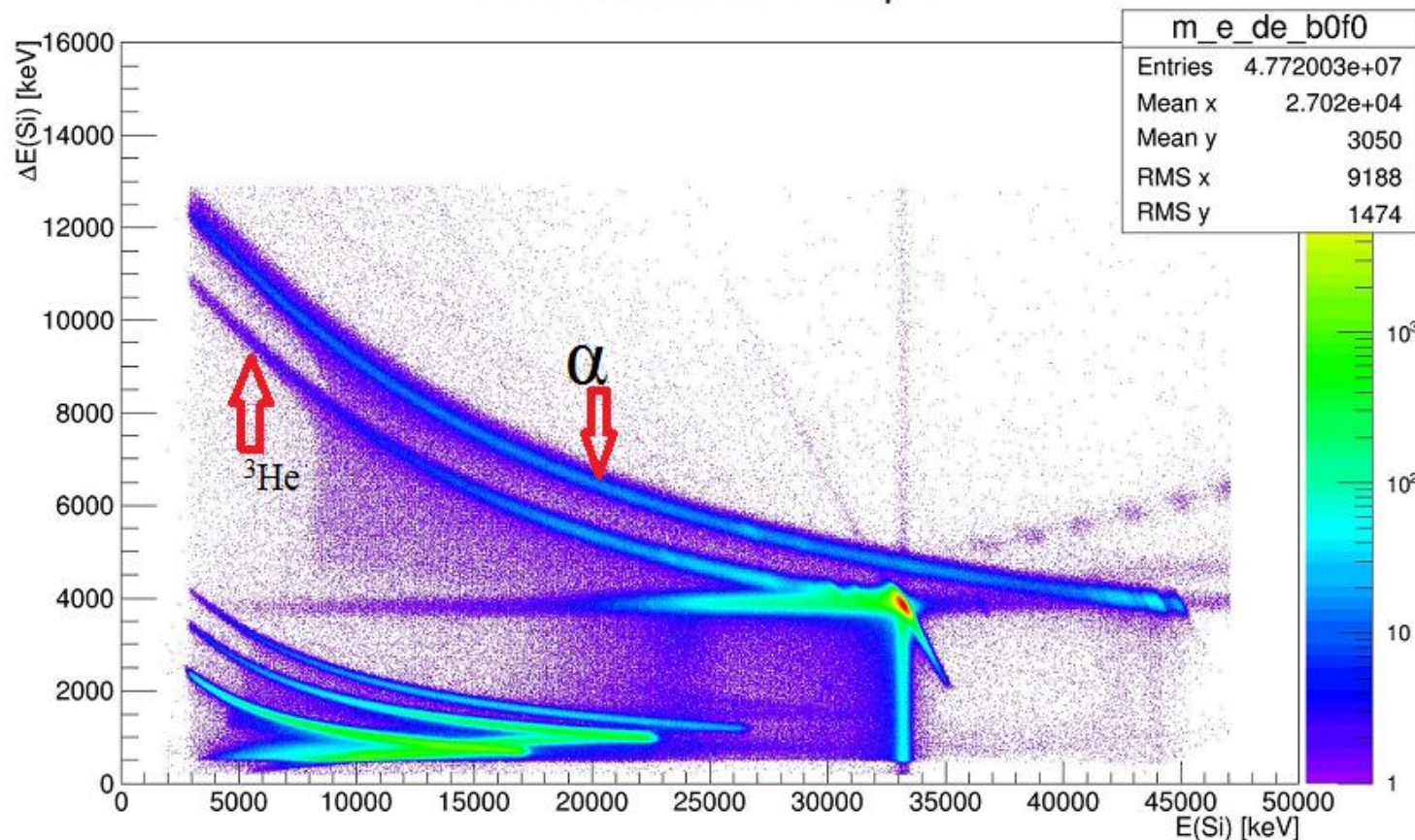
- 2.5 mg/cm^2 NatLa (99.9% ^{139}La) targets provided by iTL. (N.Y. Kheswa): 5 samples shipped to Oslo (only 3 survived)
- Elemental La oxidizes easily!
- First transportation in a vacuum-tight container (2 samples) was unsuccessful
- The foil used during the measurement was instead transported in Argon and silica gel (which absorbs moisture)
- The scattering chamber was filled with Argon during the target mounting operation



Analysis

Ungated E vs E for particle identification

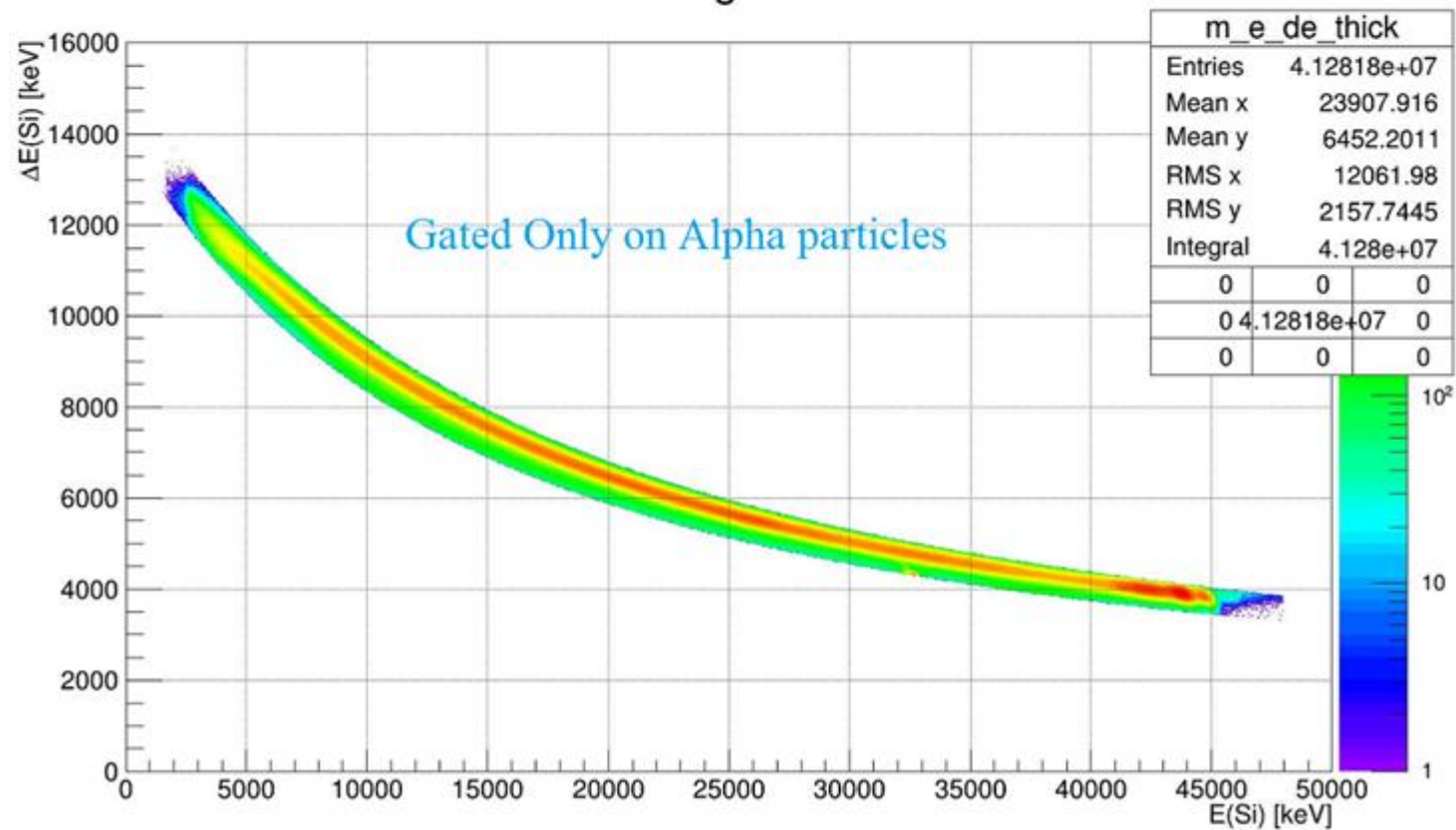
ΔE : E detector 0 strip 0



Analysis

Gated E: E Plot

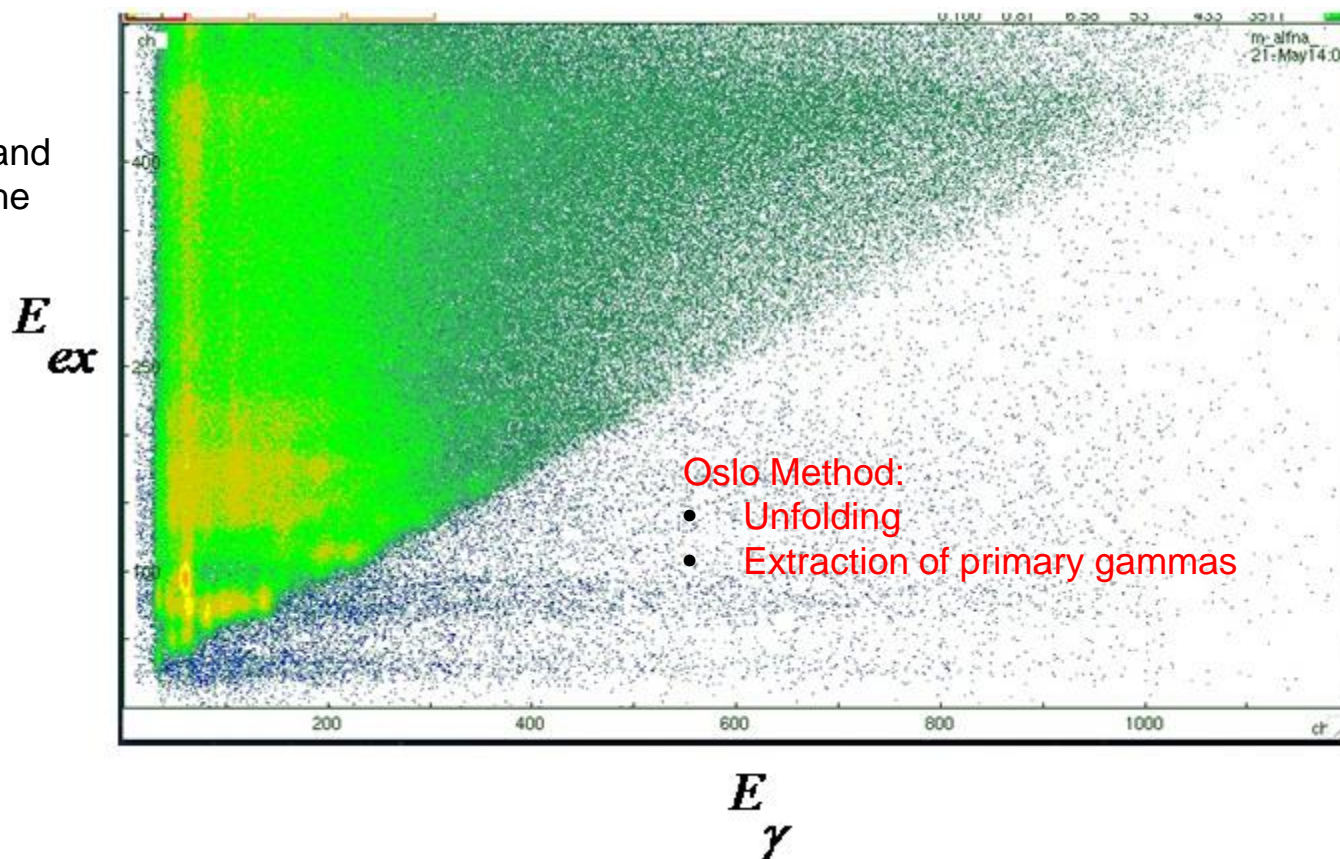
ΔE : E for all detectors together



Analysis

Particle - coincidence matrix

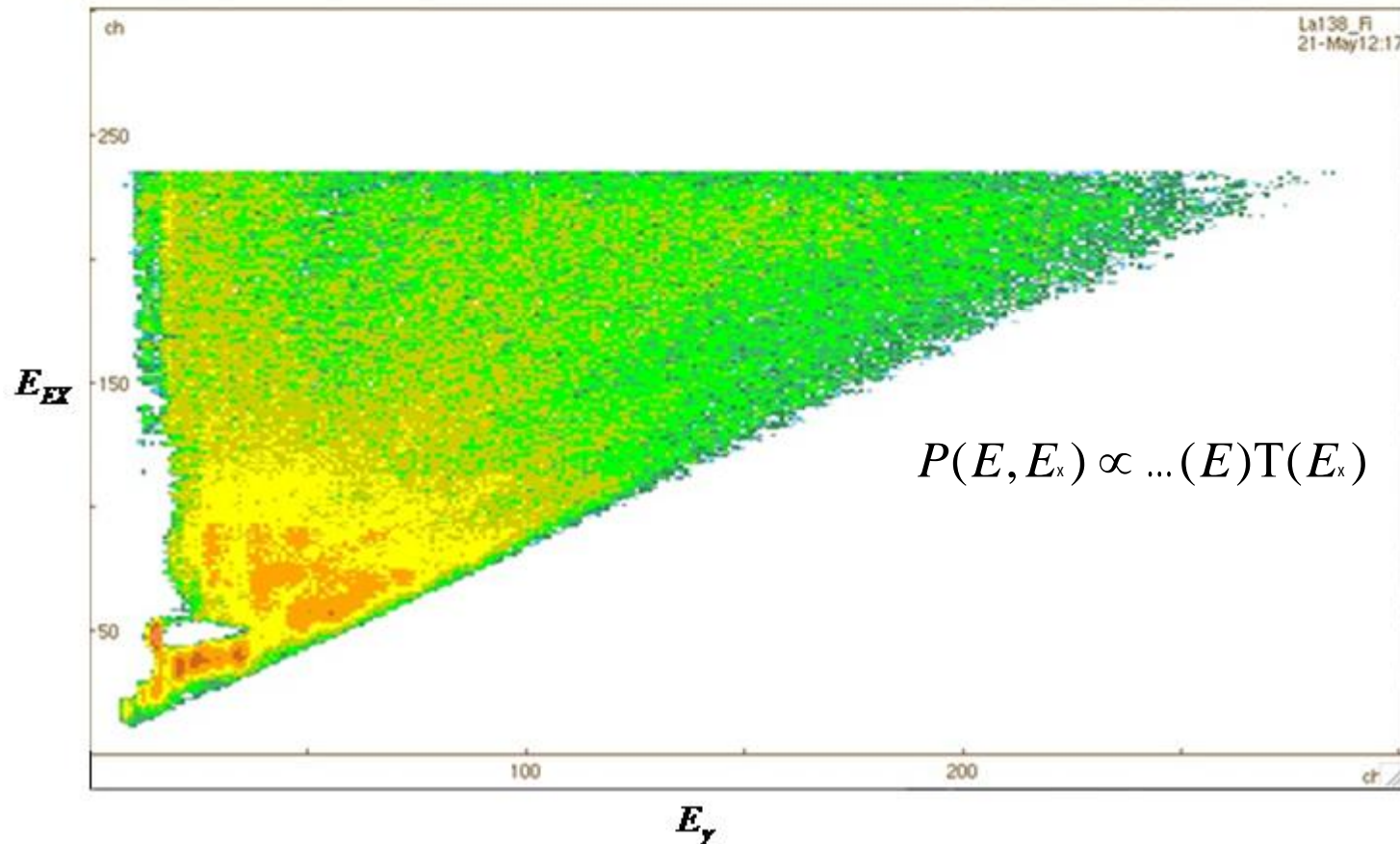
- From Q-value and kinematics of the reaction



- Gammas in coincidence with particles

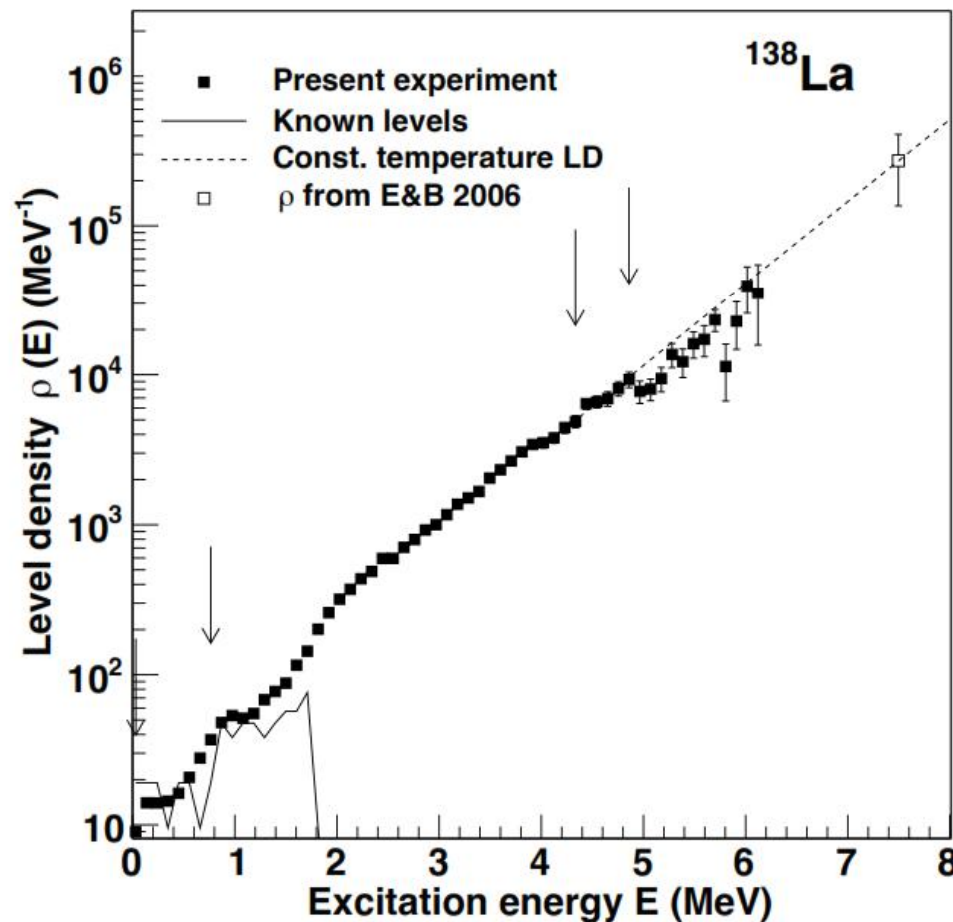
Analysis

First Generation Matrix



Preliminary Results

Level Density of La138



- (S_n) calculated using BSFG model:

$$\dots(S_n) = \frac{\exp[2\sqrt{a}(S_n - E_1)]}{12\sqrt{2\pi} * a^{1/2} (S_n - E_1)^{5/4}}$$

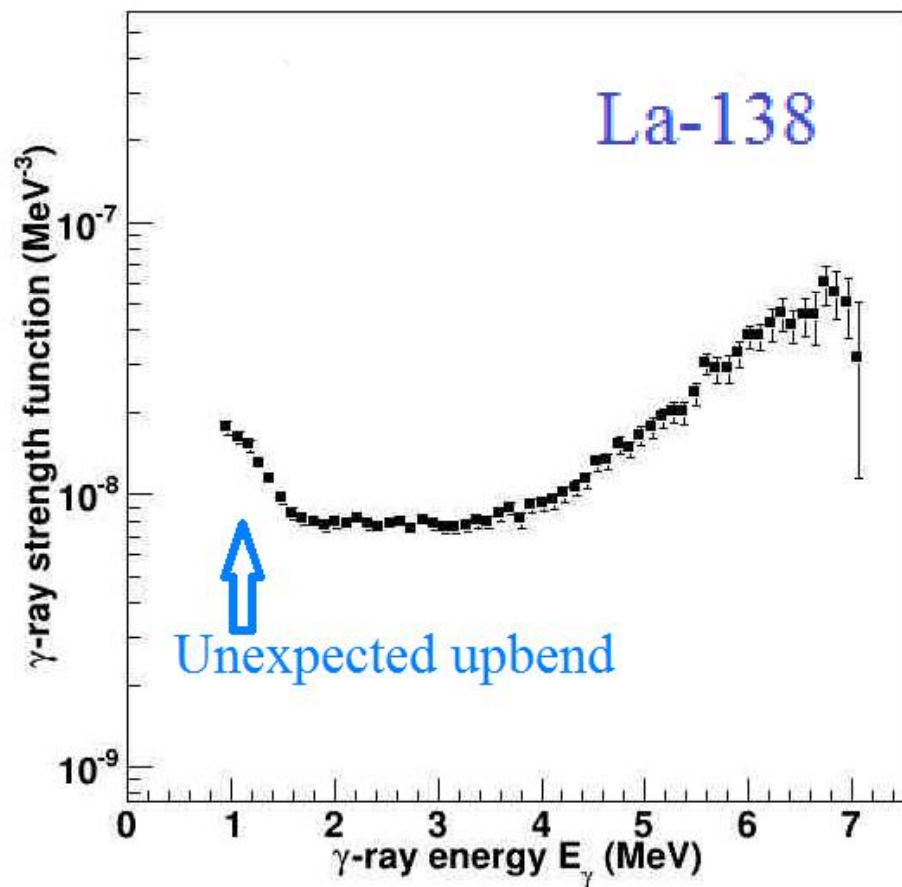
and

$$\pi^2 = 0.0146A^{5/3} \frac{1 + \sqrt{1 + 4a(U - E_1)}}{2a}$$

T. von Egidy and D. Burcurescu, PRC 72, 044311(2005)

Preliminary Results

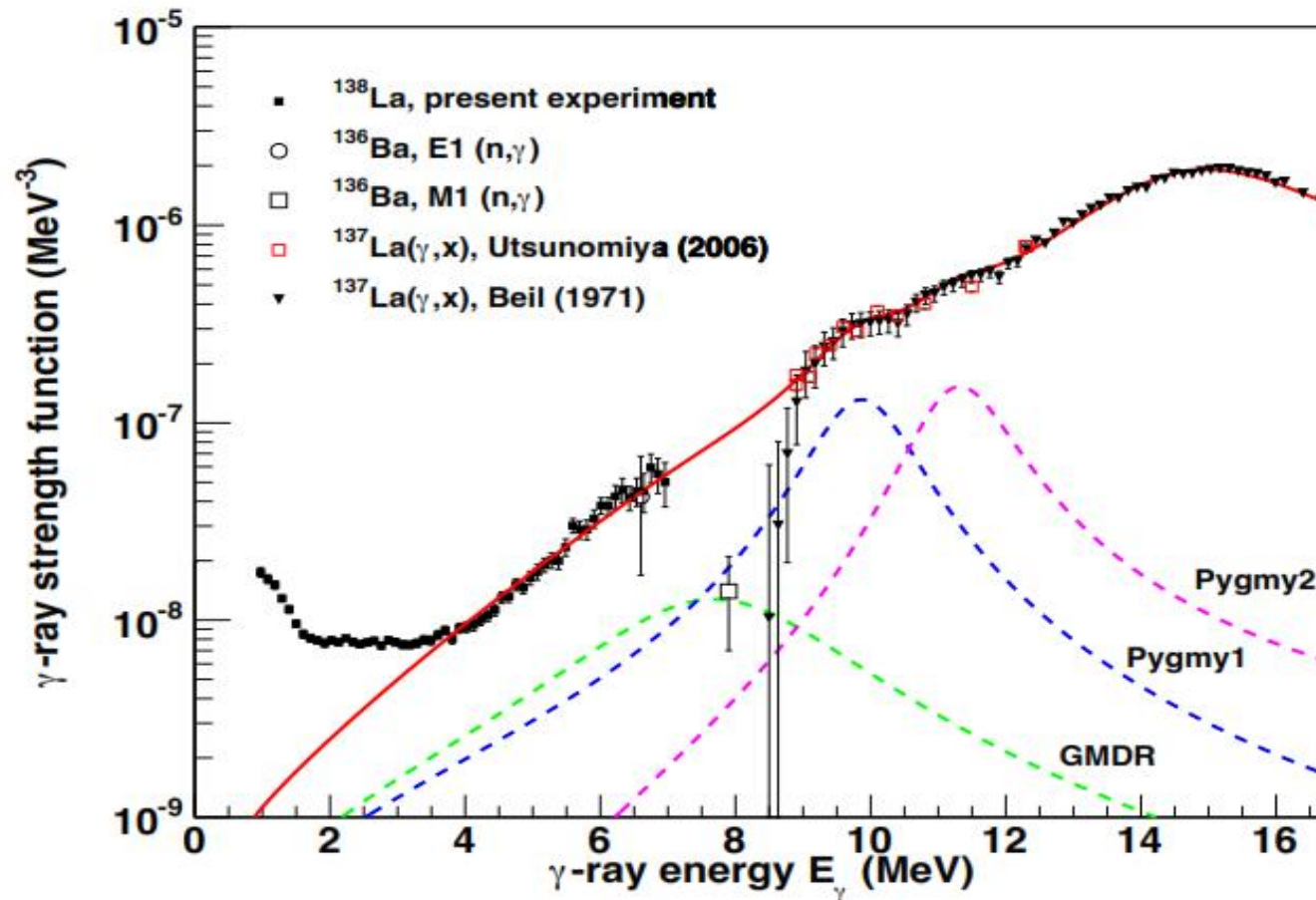
Photon Strength Function of ^{138}La



- Pronounced low energy enhancement
- ***First observation of the low-energy enhancement above mass $A \sim 106$!!!!!!***

Preliminary Results

Comparison with GDR Model and Experimental data



Clearly there is a low energy enhancement !!!!!!!!!!!!!

Conclusions and outlook

- To better understand the natural abundance of the “p-nuclei” and in particular ^{138}La , more precise production rates have to be determined: Experimental data on PSF and LD as input parameters are needed.
- A $^{139}\text{La}(^3\text{He}, \gamma)^{138}\text{La}$ experiment has been successfully performed in February 2013 at OCL.
- Preliminary results on PSF and LD have been extracted: **an unexpected up-bend at low energy in the PSF have been observed for the first time in nuclei heavier than $A \sim 106!!!$**
- Next steps: comparison with more experimental data and models.
PSF and LD for ^{139}La via the $(^3\text{He}, ^3\text{He})$ reaction.
Plans for neutron capture measurements to extract PSF via TSC analysis.

Acknowledgements



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- N.Y. Kheswa¹ for making the target
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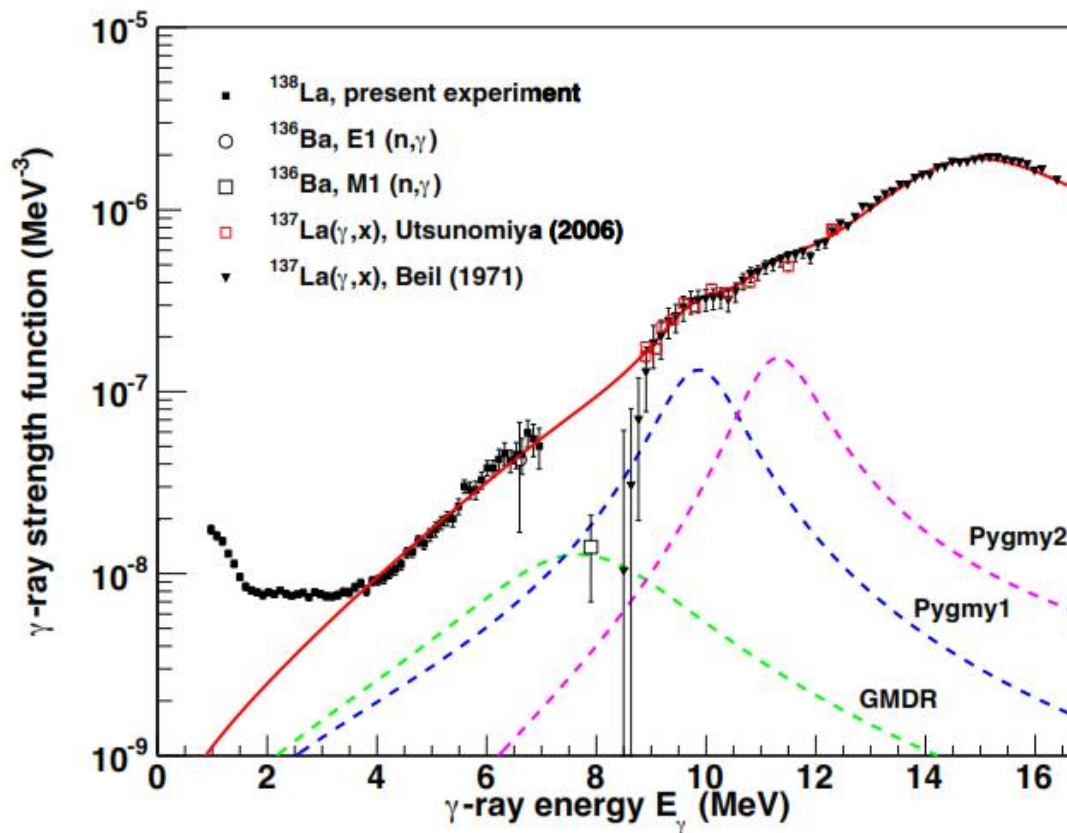
¹iThemba LABS, Somerset West, South Africa

²SAFE, University of Oslo, Oslo, Norway

Thank You All

Preliminary Results

Comparison with GDR Model and Experimental data



- *Pygmy1 and Pygmy2 are E2 and E1 resonances*
- *Models:*
 - Spin-flip Model for GMDR*
 - Isoscalar Model for Pygmy1*
 - EGLO for GEDR and Pygmy2*

Clearly there is a low energy enhancement !!!!!!!!!!!!!