



CEA-Cadarache

5TH WORKSHOP ON NUCLEAR LEVEL DENSITY AND GAMMA STRENGTH

Neutron-induced cross sections of ^{236}U , ^{238}U , ^{237}Np and ^{238}Np via the surrogate reaction method

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Outline of my talk

Motivation

I) Surrogate Method

- a) Principle
- b) Validity & Objective
- c) Recent Work
- d) Interpretation of the results
- e) Summary

II) Oslo measurements

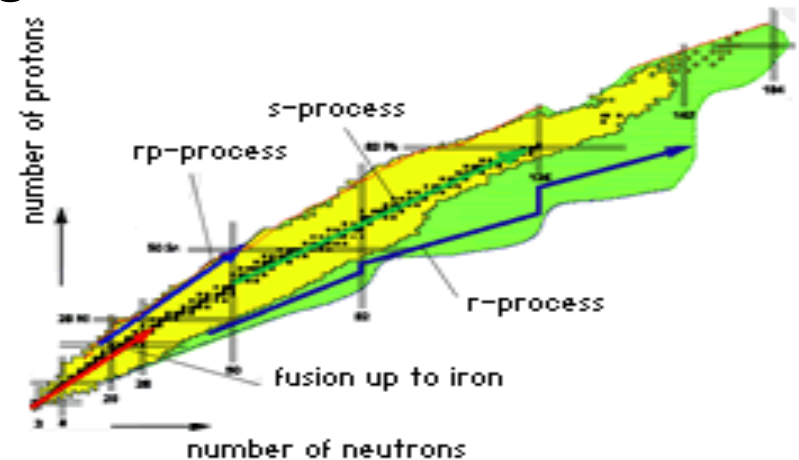
- a) Experimental set-up
- b) Investigated reactions
- c) Results
- d) Comparison with statistical model

Conclusion

Motivation

Importance of neutron-induced cross sections of short lived nuclei:

- Fundamental nuclear physics
- Reactor physics
- Stellar nucleosynthesis via r or s processes

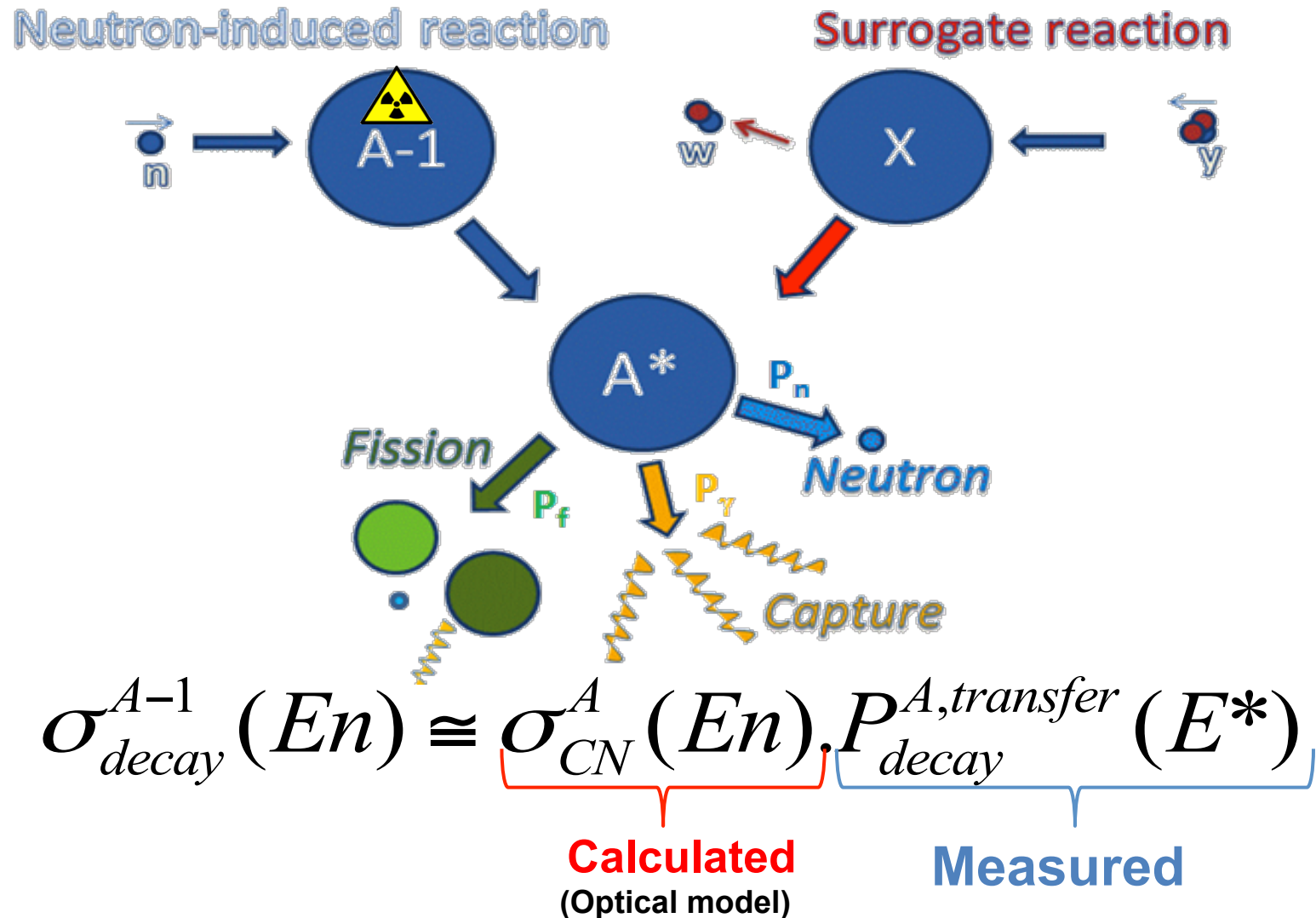


BUT these neutron-induced cross sections of short lived nuclei extremely difficult to obtain due to the radioactivity of the targets involved.



Surrogate reaction

Surrogate Experiment : principle



J.D. Cramer et H.C. Britt, Nucl. Sci. And Eng. 41 (1970) 177

Surrogate Experiment : validity & objective

$$\sigma_{decay}^{A-1}(En) \cong \underbrace{\sigma_{CN}^A(En)}_{\text{Calculated (Optical model)}} \underbrace{P_{decay}^{A,transfer}(E^*)}_{\text{Measured}}$$

Surrogate reaction only valid if :

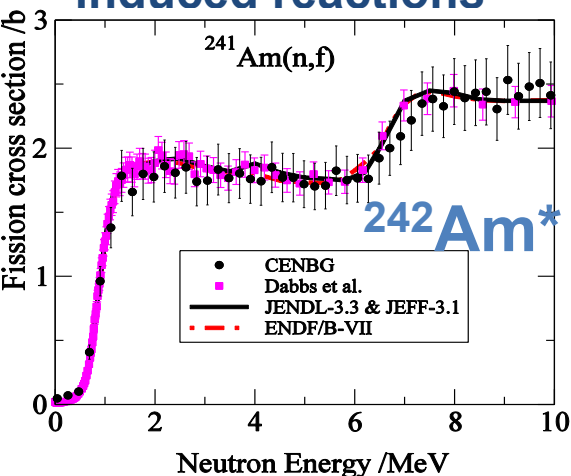
- Compound nucleus formation
- $P_{decay}^{n\text{-induced}} \approx P_{decay}^{surrogate} \rightarrow$ Similar J^π distribution or no dependence on J^π (Weisskopf-Ewing limit)

Objective :

- Validity of surrogate method by comparing surrogate data to known n-induced data.

Surrogate Experiment : Recent work

Comparison surrogate/neutron induced reactions

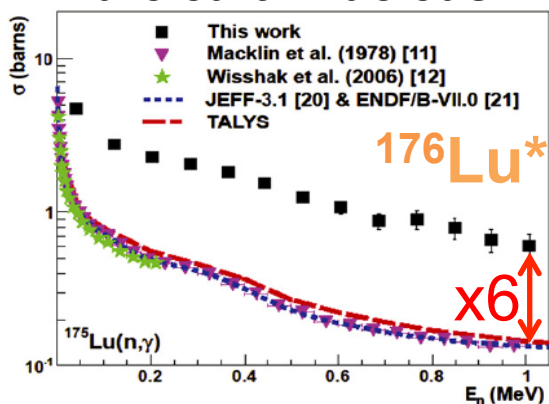


G. Kessedjian, et al., Phys. Lett B 692 (2010) 297

Fission
OK !

Example

Rare earth nucleus



G. Boutoux, et al., Phys. Lett B 712 (2012) 319

Capture
NOT OK !

Example

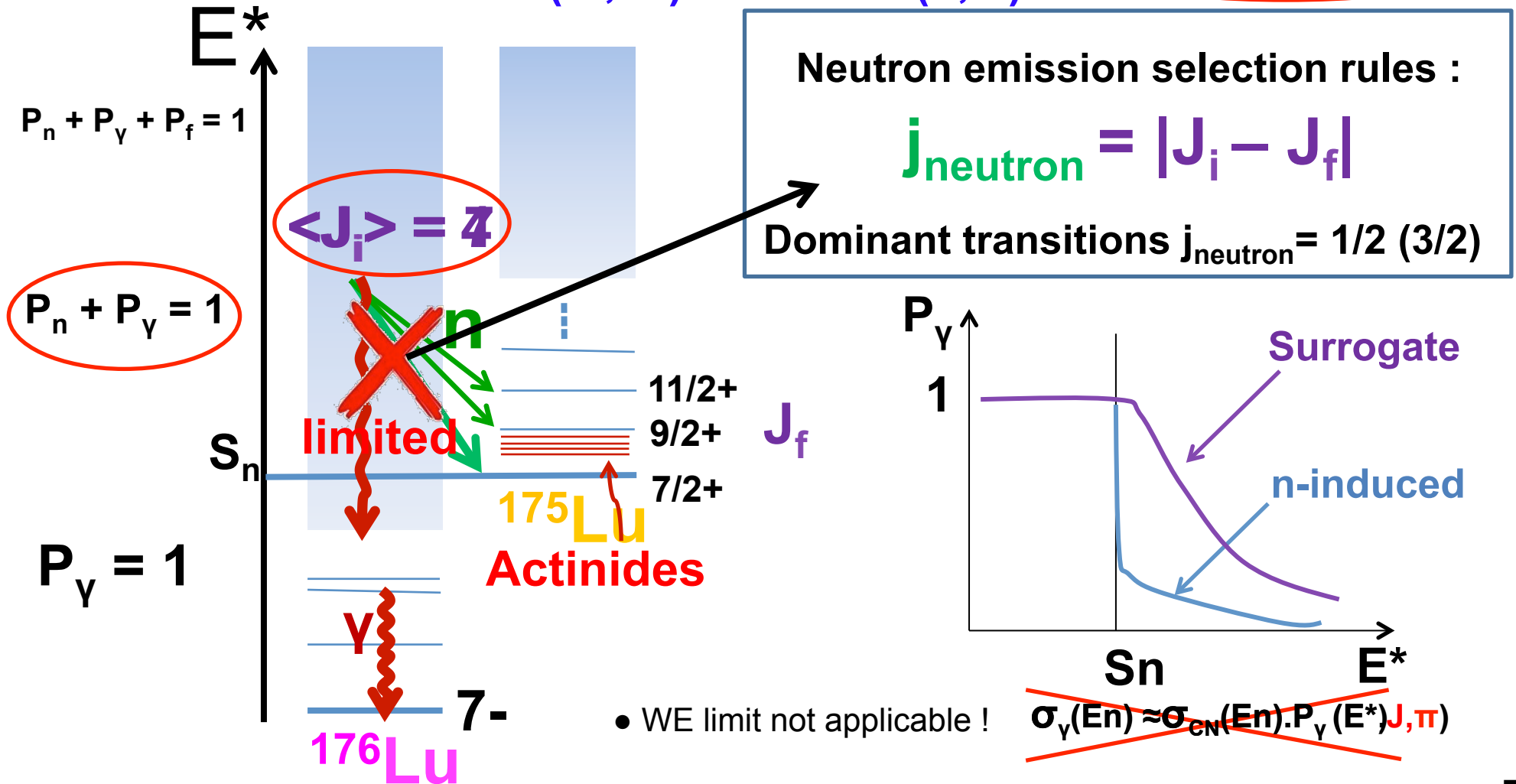
Examples of surrogate reactions performed since 2004

Desired reaction	E_n range (MeV)	Surrogate reaction	Type	Reference
(n, f) cross sections				
$^{230}\text{Th}(n, f)$	0.5–10	$^{232}\text{Th}(^3\text{He}, \alpha)$	absolute	Petit et al. (2004)
$^{230}\text{Th}(n, f)$	0.22–25	$^{232}\text{Th}(^3\text{He}, \alpha)$	ratio	Goldblum et al. (2009)
$^{231}\text{Th}(n, f)$	0.36–25	$^{232}\text{Th}(^3\text{He}, ^3\text{He}')$	ratio	Goldblum et al. (2009)
$^{231}\text{Pa}(n, f)$	0.5–10	$^{232}\text{Th}(^3\text{He}, t)$	absolute	Petit et al. (2004)
$^{233}\text{Pa}(n, f)$	0.5–10	$^{232}\text{Th}(^3\text{He}, p)$	absolute	Petit et al. (2004)
$^{233}\text{Pa}(n, f)$	11.5–16.5	$^{232}\text{Th}(^6\text{Li}, \alpha)$	ratio	Nayak et al. (2008)
$^{233}\text{U}(n, f)$	0.4–18	$^{234}\text{U}(\alpha, \alpha')$	ratio	Leshner et al. (2009)
$^{236}\text{U}(n, f)$	0–20	$^{238}\text{U}(^3\text{He}, \alpha)$	absolute, ratio	Lyles et al. (2007a)
$^{237}\text{U}(n, f)$	0–13	$^{238}\text{U}(d, d')$	ratio	Plettner et al. (2005)
$^{237}\text{U}(n, f)$	0–20	$^{238}\text{U}(\alpha, \alpha')$	ratio	Burke et al. (2006)
$^{239}\text{U}(n, f)$	0–20	$^{238}\text{U}(^{18}\text{O}, ^{16}\text{O})$	ratio	Burke et al. (2011)
$^{237}\text{Np}(n, f)$	10–20	$^{238}\text{U}(^3\text{He}, t)$	absolute, ratio	Basunia et al. (2009)
$^{238}\text{Pu}(n, f)$	0–20	$^{239}\text{Pu}(\alpha, \alpha')$	ratio	Ressler et al. (2011)
$^{241}\text{Am}(n, f)$	0–10	$^{243}\text{Am}(^3\text{He}, \alpha)$	absolute	Kessedjian et al. (2010)
$^{242}\text{Cm}(n, f)$	0–10	$^{243}\text{Am}(^3\text{He}, t)$	absolute	Kessedjian et al. (2010)
$^{243}\text{Cm}(n, f)$	0–3	$^{243}\text{Am}(^3\text{He}, d)$	absolute	Kessedjian et al. (2010)
(n, γ) cross sections				
$^{155}\text{Gd}(n, \gamma)$	0.05–3.0	$^{156}\text{Gd}(p, p')$	absolute, ratio	Scielzo et al. (2010)
$^{157}\text{Gd}(n, \gamma)$	0.05–3.0	$^{158}\text{Gd}(p, p')$	absolute, ratio	Scielzo et al. (2010)
$^{161}\text{Dy}(n, \gamma)$	0.13–0.56	$^{162}\text{Dy}(^3\text{He}, ^3\text{He}')$	ratio	Goldblum et al. (2010)
$^{170}\text{Yb}(n, \gamma)$	0.165–0.405	$^{171}\text{Yb}(^3\text{He}, ^3\text{He}')$	ratio	Goldblum et al. (2008)
$^{170}\text{Yb}(n, \gamma)$	0.225–0.465	$^{172}\text{Yb}(^3\text{He}, \alpha)$	ratio	Goldblum et al. (2008)
$^{171}\text{Yb}(n, \gamma)$	0.12–0.24	$^{171}\text{Yb}(d, p)$	ratio	Hatarik et al. (2010)
$^{233}\text{Pa}(n, \gamma)$	0–1	$^{232}\text{Th}(^3\text{He}, p)$	absolute	Boyer et al. (2006)
$^{235}\text{U}(n, \gamma)$	0.9–3.3	$^{235}\text{U}(d, p)$	ratio	Allmond et al. (2009)
$^{237}\text{U}(n, \gamma)$	0.2–1.0	$^{238}\text{U}(\alpha, \alpha')$	absolute, ratio	Bernstein et al. (2006);
$^{232}\text{Th}(n, \gamma)$	0–1.2	$^{232}\text{Th}(d, p)$	absolute	J. Wilson et al. (2012)
$^{175}\text{Lu}(n, \gamma)$	0–1	$^{174}\text{Yb}(^3\text{He}, p)$	absolute	G. Boutoux et al. (2012)
$^{172}\text{Yb}(n, \gamma)$	0–1	$^{174}\text{Yb}(^3\text{He}, \alpha)$	absolute	G. Boutoux et al. (2012)

Jutta Escher, et al., Rev. Mod. Phys. 84 (2012) 353

Surrogate Experiment : Interpretation of the results

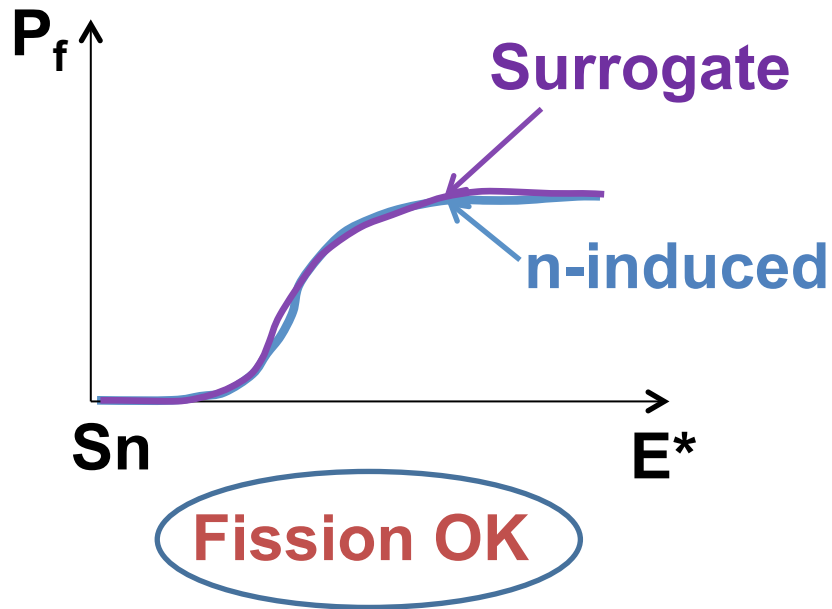
Why $\sigma_y(\text{surrogate}) \gg \sigma_y(\text{n-induced})$



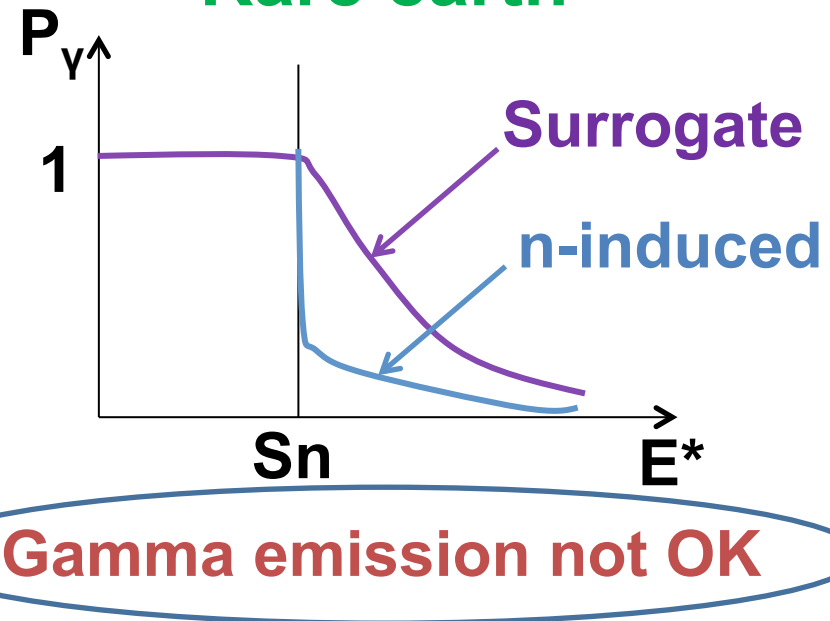
Surrogate Experiment : Summary

Schematical view of the results of the probabilities

Actinides



Rare earth

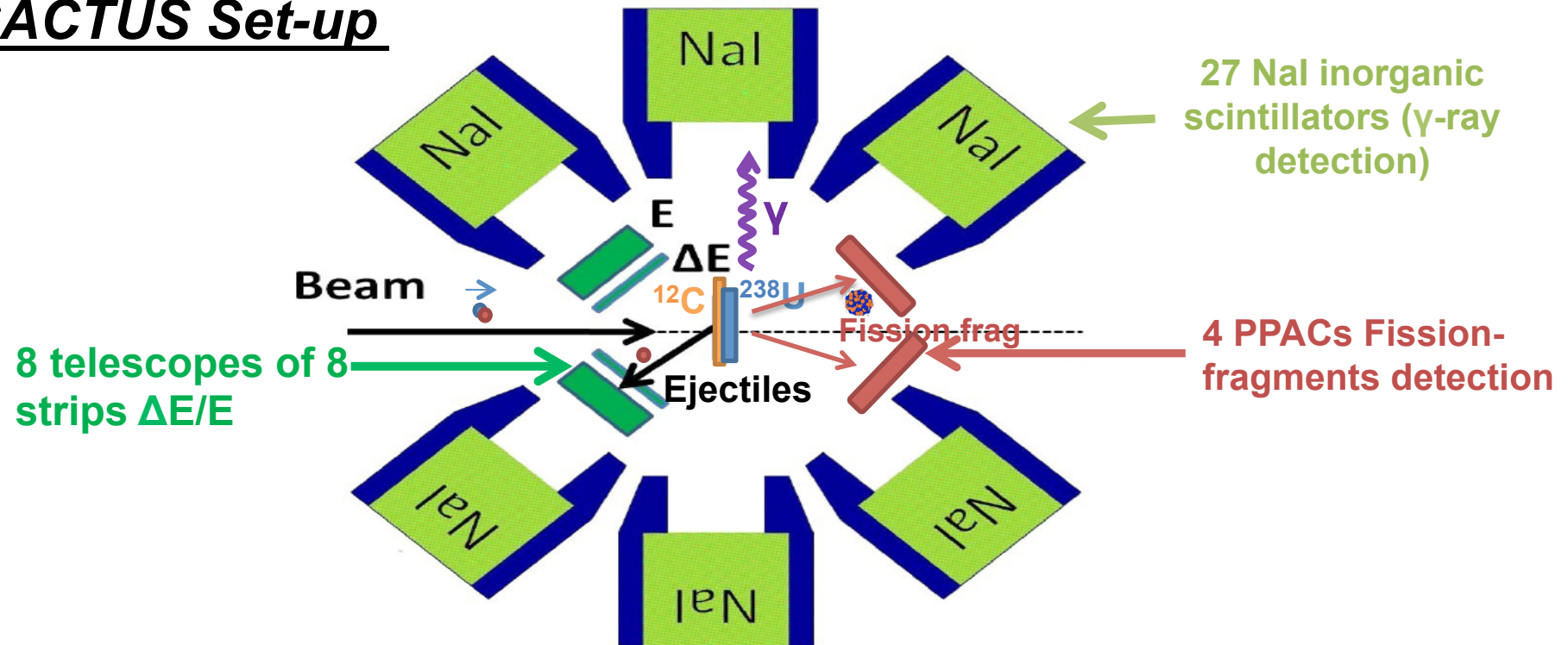


What about P_f AND P_γ for Actinides ?

Aim : Check if fission AND gamma decay probabilities agree with neutron data by measuring both **simultaneously in the actinides region**

Oslo measurements: Experimental set-up

CACTUS Set-up



$$P_{f,\gamma} = \frac{N_{f,\gamma}^{\text{coinc}}}{N_{\text{ej}} \cdot \epsilon_{f,\gamma}}$$

- ❖ High detection efficiency
- ❖ Measurement of Fission & gamma-decay probabilities

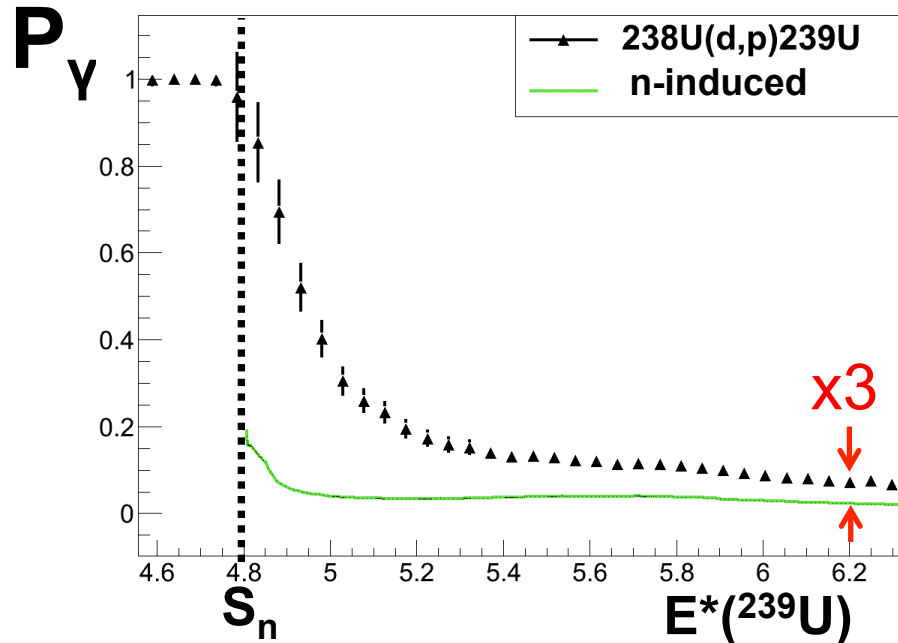
Oslo measurements: Investigated reactions

Neutron- induced reaction	Corresponding surrogate reaction	Quantity measured
$n + {}^{238}\text{U}$ ($T_{1/2} = 4\text{Gy}$)	${}^{238}\text{U} + d \rightarrow {}^{239}\text{U}^* + p$	$P_f + P_\gamma$
$n + {}^{236}\text{U}$ ($T_{1/2} = 23\text{My}$)	${}^{238}\text{U} + {}^3\text{He} \rightarrow {}^{237}\text{U}^* + {}^4\text{He}$	$P_f + P_\gamma$
$n + {}^{237}\text{Np}$ ($T_{1/2} = 2\text{My}$)	${}^{238}\text{U} + {}^3\text{He} \rightarrow {}^{238}\text{Np}^* + t$	$P_f + P_\gamma$
$n + {}^{238}\text{Np}$ ($T_{1/2} = 2\text{ days}$)	${}^{238}\text{U} + {}^3\text{He} \rightarrow {}^{239}\text{Np}^* + d$	$P_f + P_\gamma$

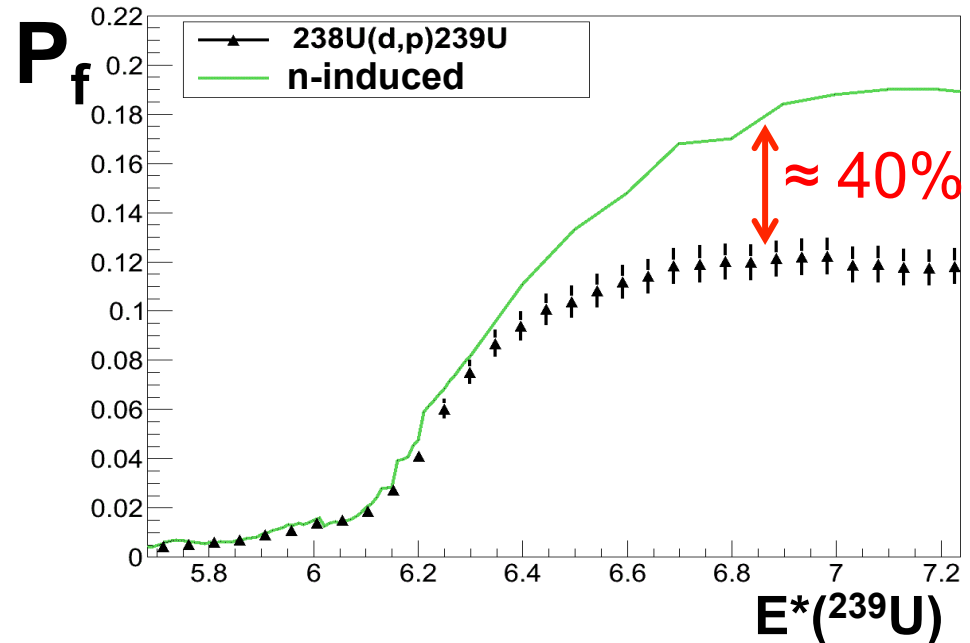
$$\sigma_{decay}^{A-1}(En) \cong \underbrace{\sigma_{CN}^A(En)}_{\text{Calculated (Optical model)}} \underbrace{P_{decay}^{A,transfer}(E^*)}_{\text{Measured}}$$

Oslo measurements: Results 1/3

(neutron-induced) $^{238}\text{U} + n \rightarrow ^{239}\text{U}^* \leftarrow ^{238}\text{U} (d,p) \text{ (surrogate)}$

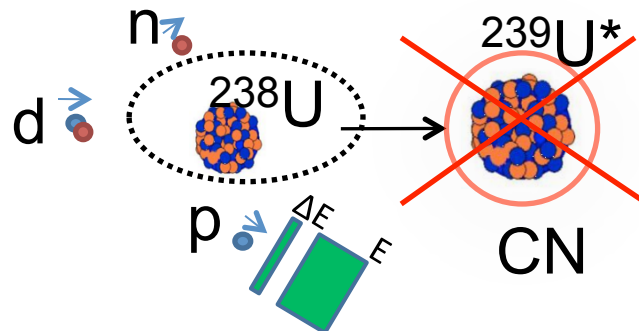


P_γ : Large discrepancies between surrogate/neutron-induced data



P_f : Discrepancies with neutron data above fission threshold

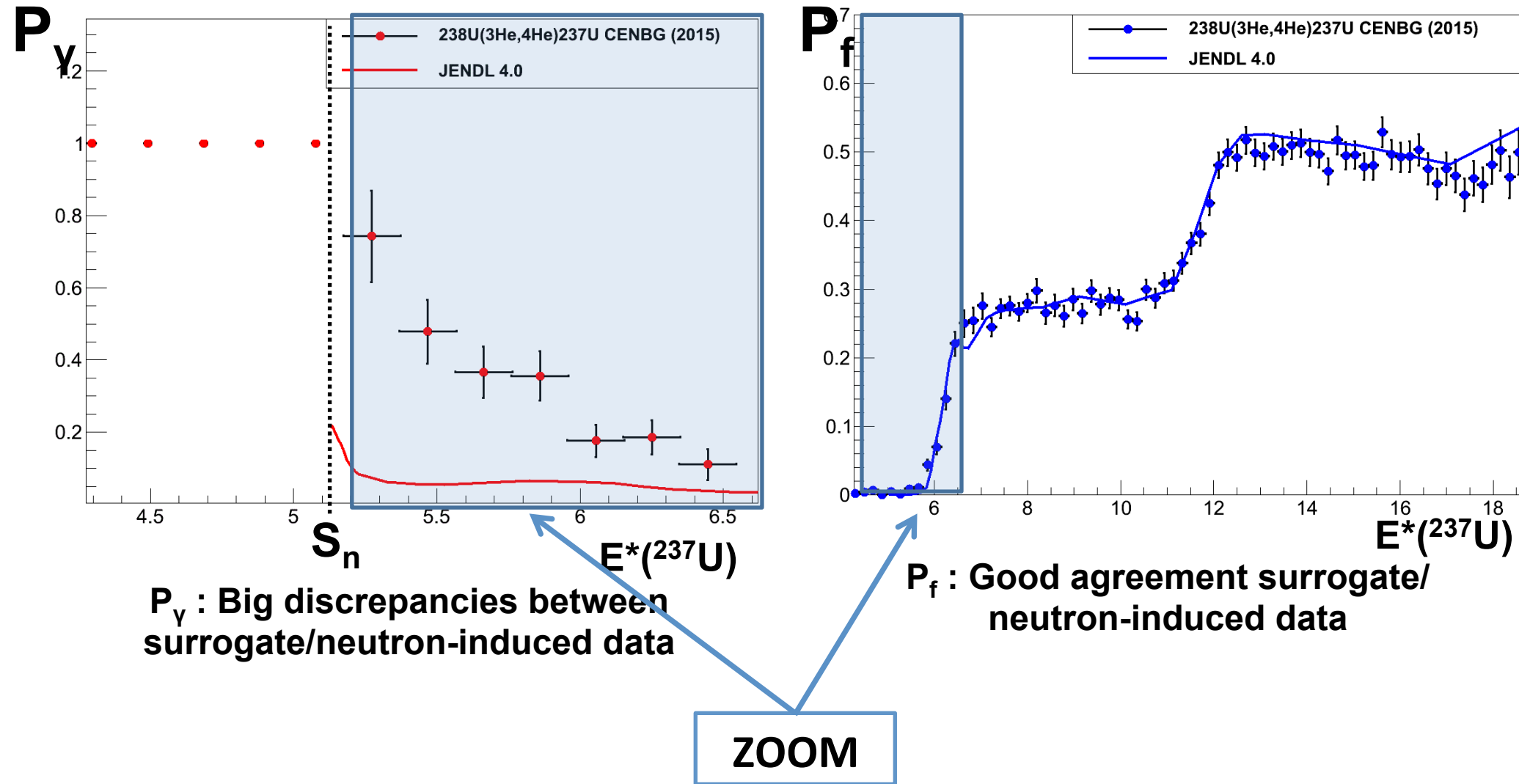
Deuteron Break-up (illustration)



$$P_{f,\gamma} = \frac{N_{f,\gamma}^{\text{coinc}}}{N_{\text{éj}} \cdot \epsilon_{f,\gamma}}$$

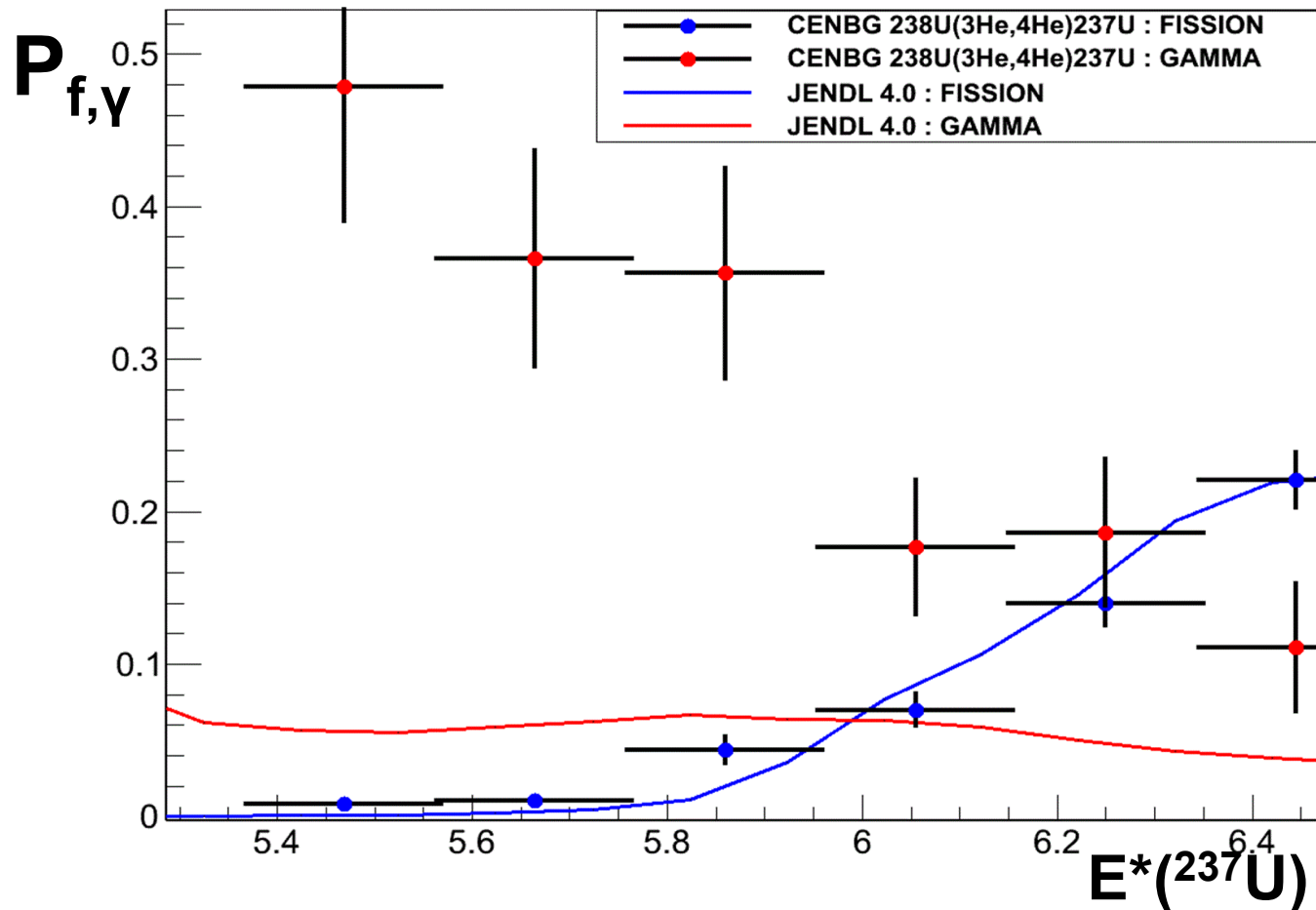
Oslo measurements: Results 2/3

(neutron-induced) $^{236}\text{U} + n \rightarrow ^{237}\text{U}^* \leftarrow ^{238}\text{U}(^3\text{He}, ^4\text{He})$ (surrogate)



Oslo measurements: Results 2/3

(neutron-induced) $^{236}\text{U} + n \rightarrow ^{237}\text{U}^* \leftarrow ^{238}\text{U}(^3\text{He}, ^4\text{He})$ (surrogate)

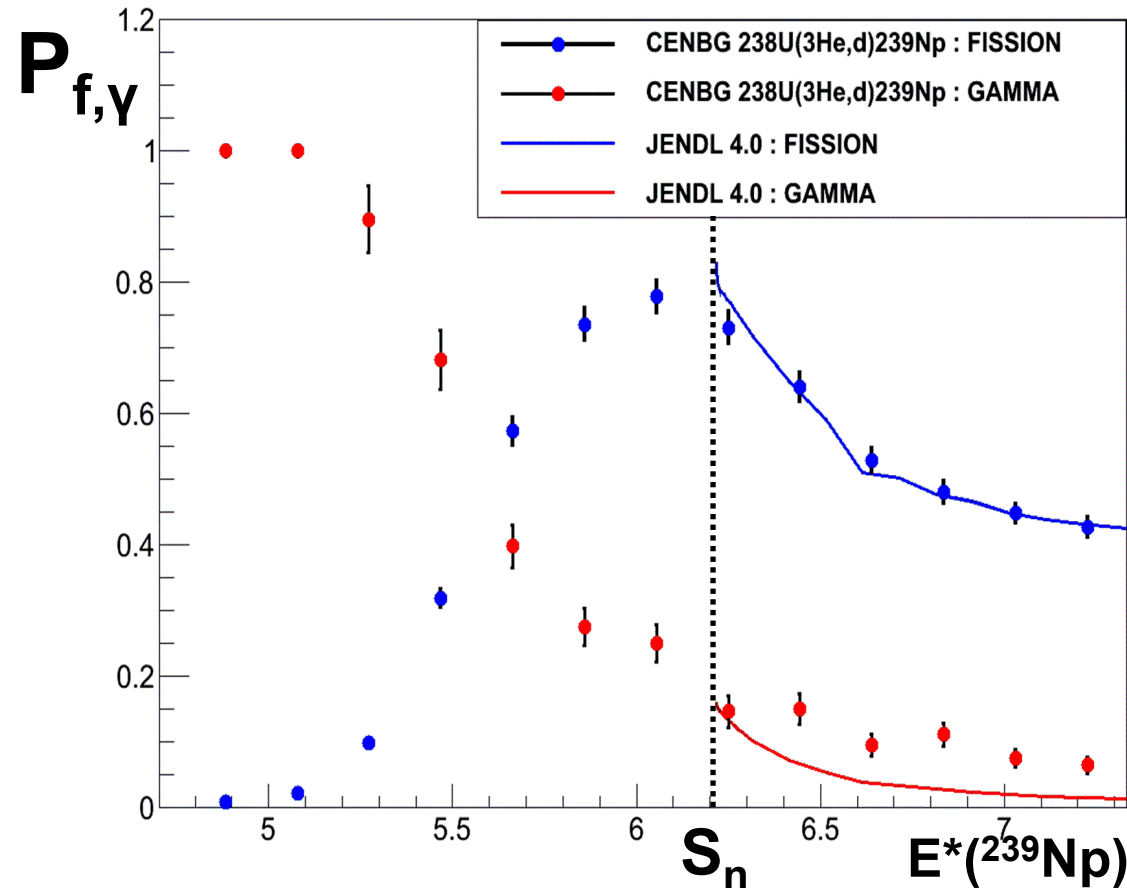


Observations :

- P_γ is much more sensitive to the spin differences than fission

Oslo measurements: Results 3/3

(neutron-induced) $^{238}\text{Np} + n \rightarrow ^{239}\text{Np}^* \leftarrow ^{238}\text{U}(^3\text{He},d)$ (surrogate)



Observations :

- P_γ is much more sensitive to the spin differences than fission
- Can we explain these results with statistical model calculations ?

Oslo measurements: Comparison with statistical models 1/2

Statistical model (Hauser-Feshbach) approach :

EVITA code (CEA France)

Aim : Use of statistical model to see if it reproduces/can explain the different sensibilities to the spin parities distribution for the gamma/fission decay

STEP 1 : Fix the parameters of the model that reproduce the neutron data (LD, γ SF...)

$$P_{\text{decay}} = f(\text{LD}, \gamma\text{SF}, \dots)$$

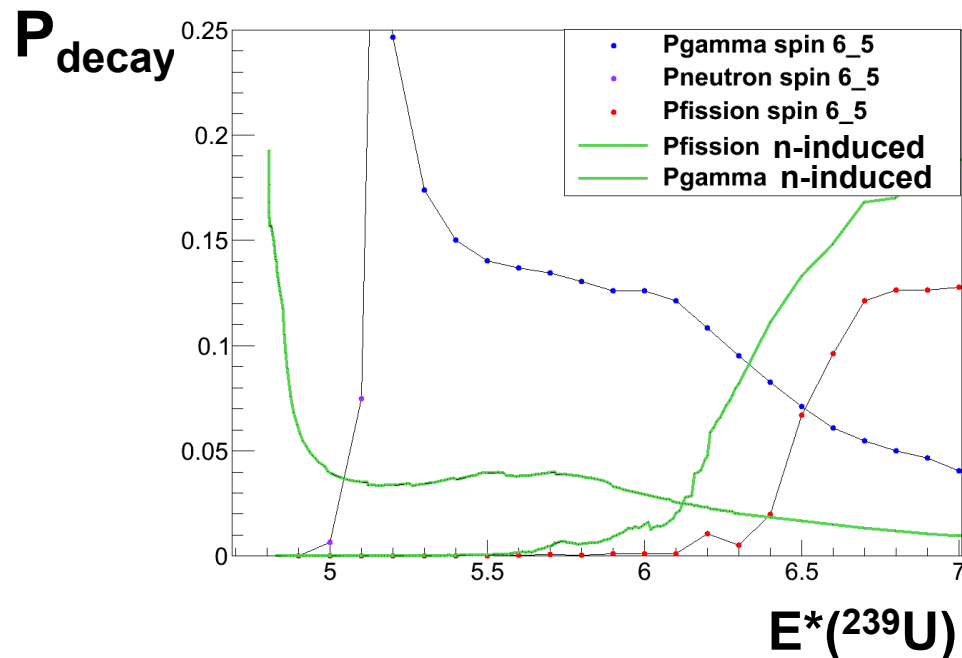
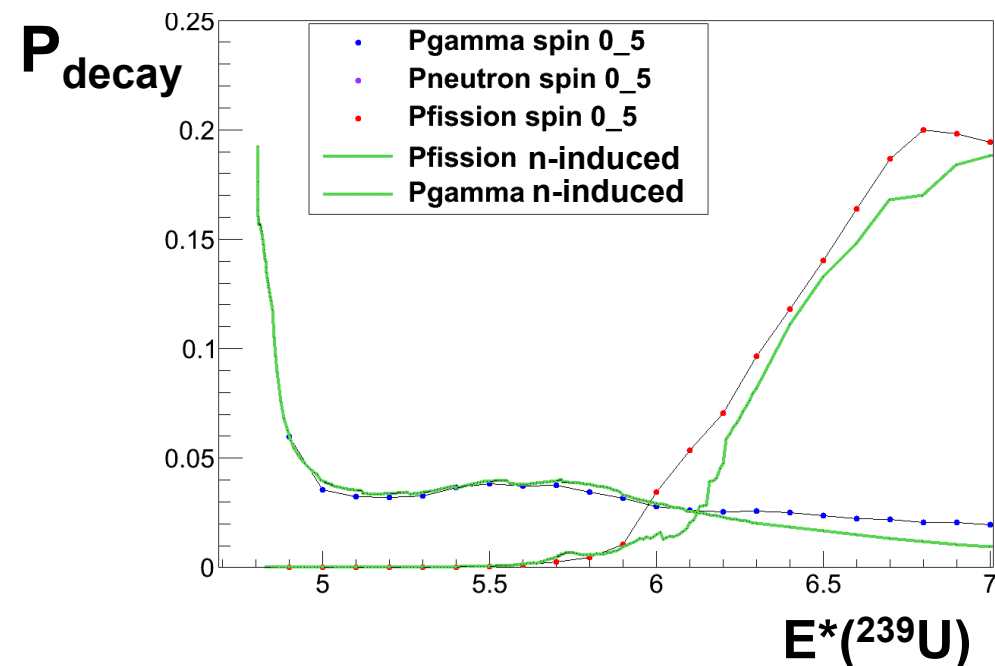
STEP 2 : Use the sensibility to spin of P_{γ} to extract information on the populated spin distribution $\langle J^{\text{surrogate}} \rangle$ of the transfer reaction

STEP 3 : Implementation of $\langle J^{\text{surrogate}} \rangle$ in the code to extract the calculated fission probability $P_{\text{f}}^{\text{EVITA}}$ and compare it with the experimental one.

Oslo measurements: Comparison with statistical models 2/2

STEP 1 : Fix the parameters of the model that reproduce the neutron data (LD, γ SF...)

(neutron-induced) $^{238}\text{U} + n \rightarrow ^{239}\text{U}^* \leftarrow ^{238}\text{U} + d$ (surrogate)



Observations :

➤ P_Y and P_f are both very sensitive to the $J\pi$

➡ In contradiction with what we observe → can not reproduce surrogate data ?

➡ Surrogate data represent an important test to statistical models

Conclusion

The surrogate method

- The surrogate method is the only way to obtain information of **very radioactive nuclei** ($T_{1/2} < \text{few days}$)

Results

- Fission : **Cross section** measurements are comparable to neutron data
- Gamma emission : **Cross section** measurements are **NOT** comparable to neutron data → use it for fixing statistical model parameters

Observations

- Fission **much less sensitive** to the spin distribution than gamma emission

Statistical model calculations

- Seem not to be able to reproduce experimental surrogate data

THANK YOU FOR YOUR ATTENTION

1) Analyse de la voie $^{238}\text{U}(d,p)^{239}\text{U} \leftrightarrow n + ^{238}\text{U}$

Equation conservation en énergie

$$E^*(\text{CN}) = E_{\text{pro}} + Q_{\text{reaction}} - E_{\text{ejectile}} - E_{\text{recul}}(\text{CN}) - E_{\text{recul}}$$

Connues

Calibration

Se calcule



$E^*(\text{CN})$ = Energie d'excitation du CN formé

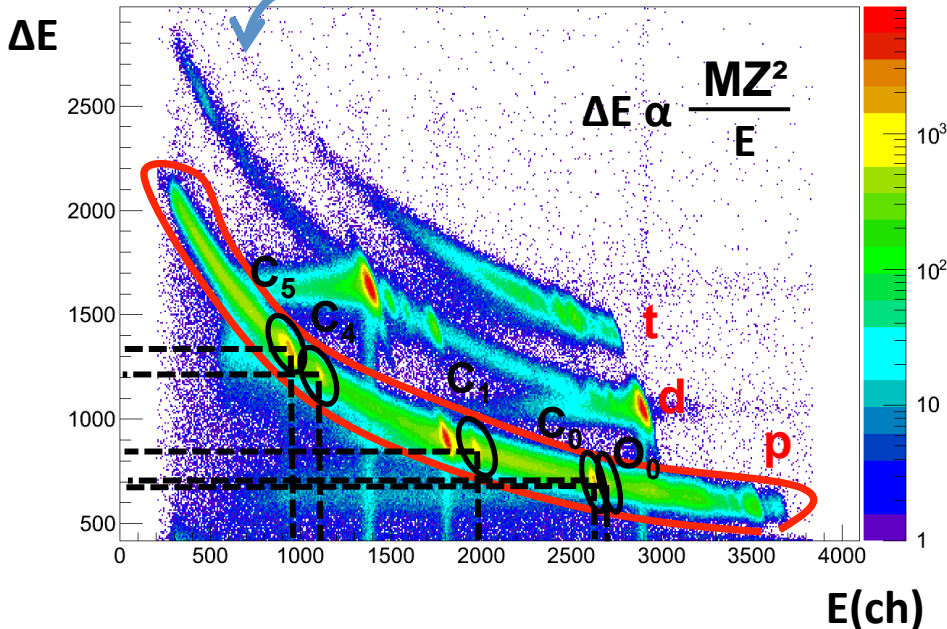
E_{pro} = Energie du faisceau (15 MeV)

$Q_{\text{reaction}} = \Delta mc^2 = 2.582 \text{ MeV}$

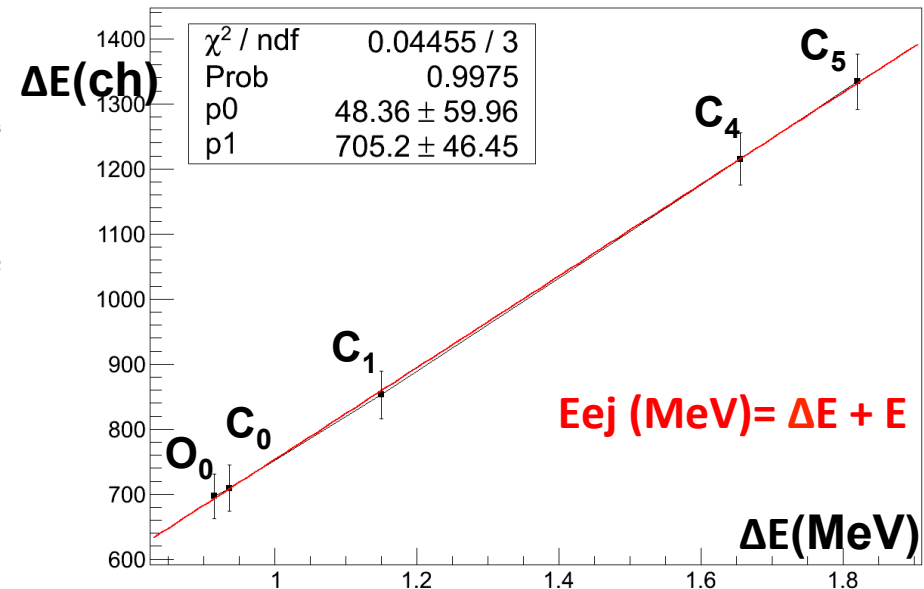
$E_{\text{ejectile}} = \Delta E + E$

$E_{\text{recul}}(\text{CN})$ = Energie de recul du CN

A) Identification du noyau composé formé



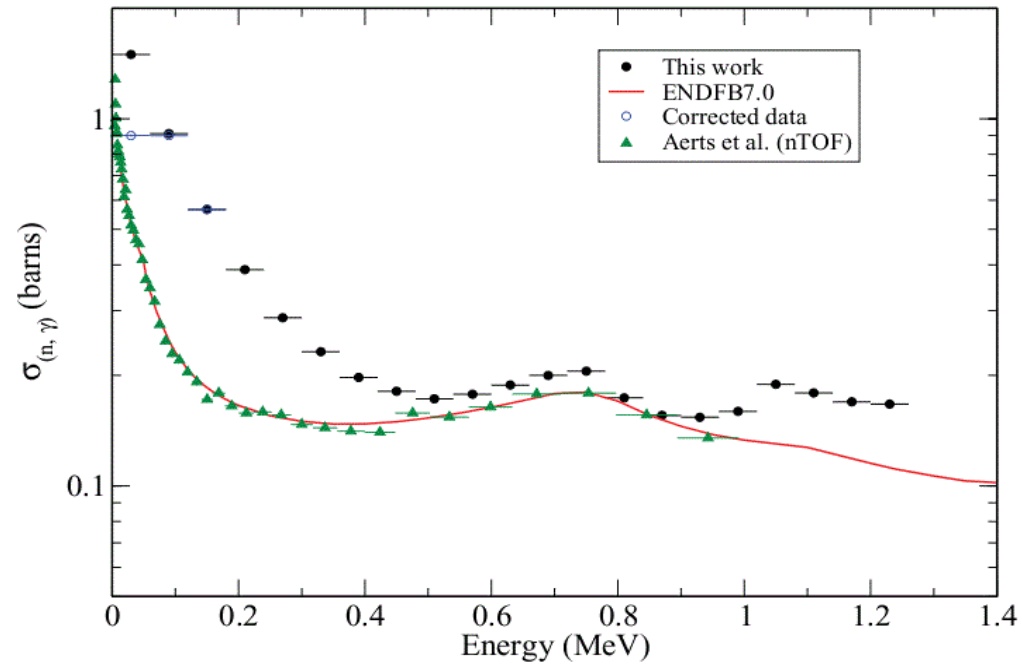
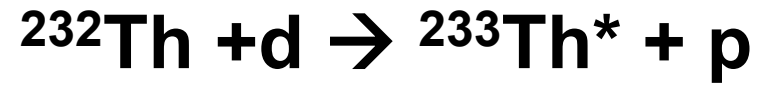
B) Calibration en énergie de la réaction



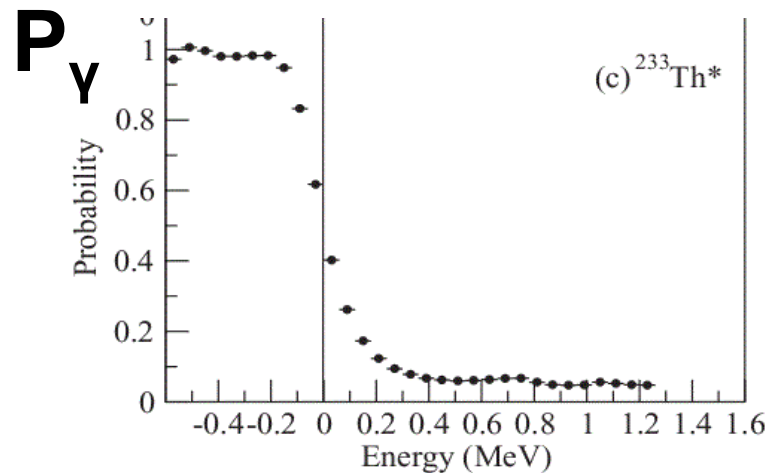
Neutron induced reaction



Surrogate reaction

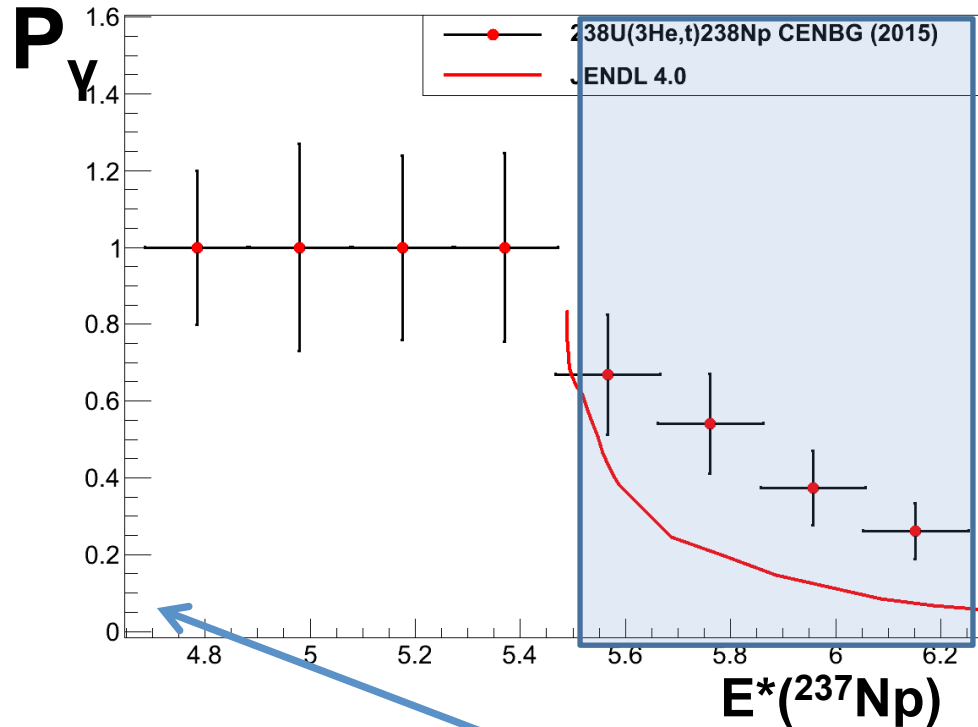


J. N. Wilson et al, Phys. Ev C 85, 034607 (2012)

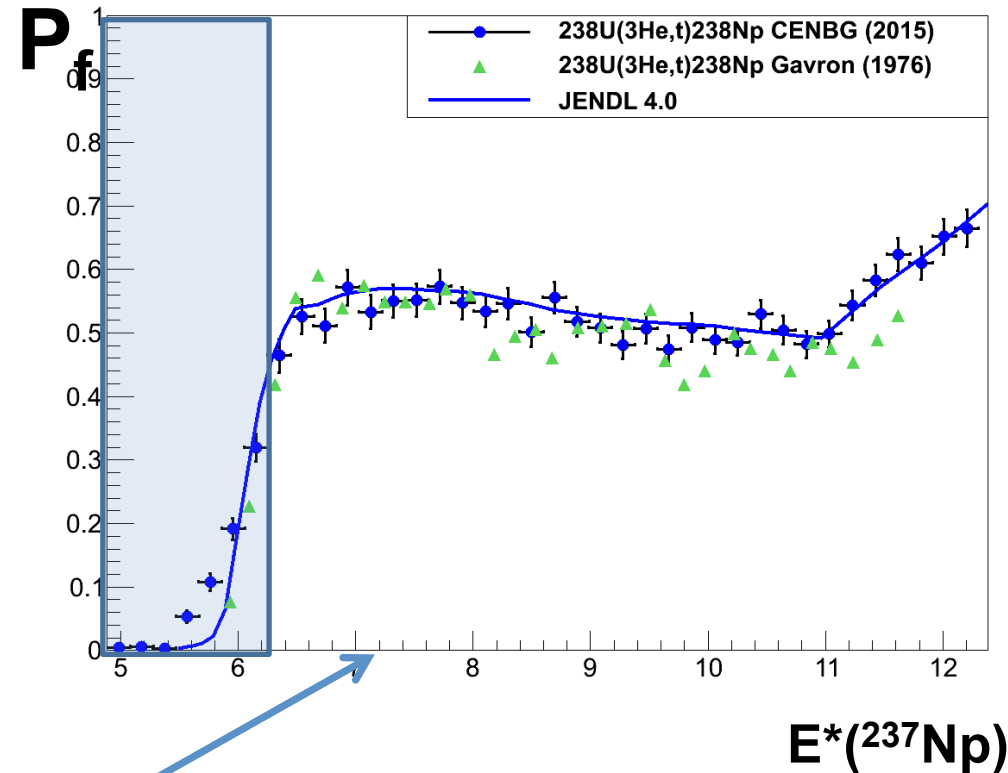


Oslo measurements: Results 1/2

(neutron-induced) $^{238}\text{U} + n \rightarrow ^{238}\text{Np}^* \leftarrow ^{237}\text{Np} + t$ (surrogate)



P_f : Good agreement surrogate/
neutron-induced data

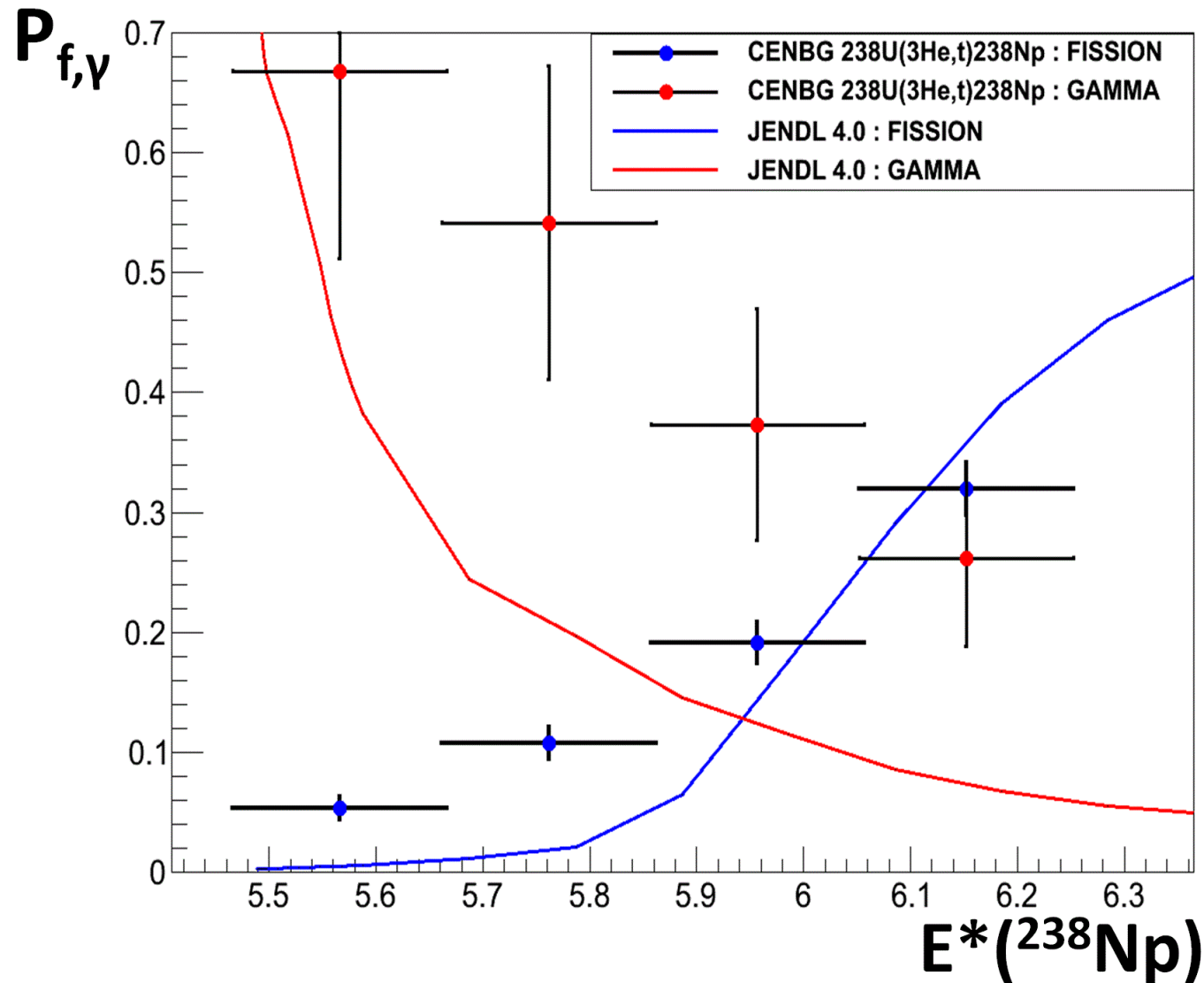


P_Y : Big discrepancies between
surrogate/neutron-induced data

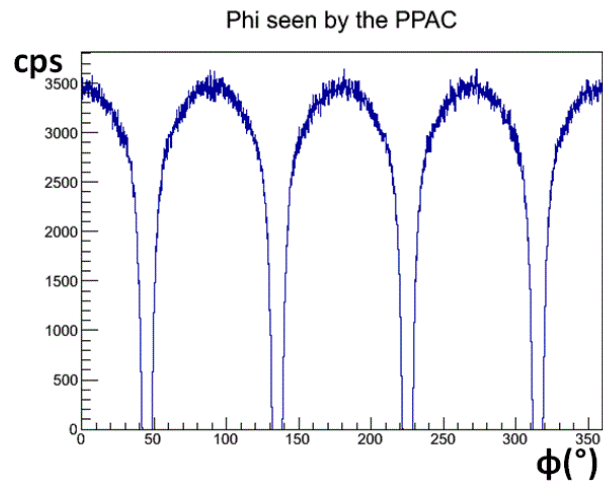
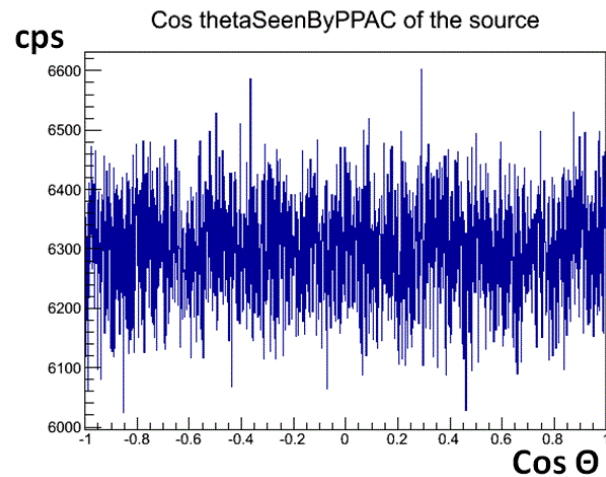
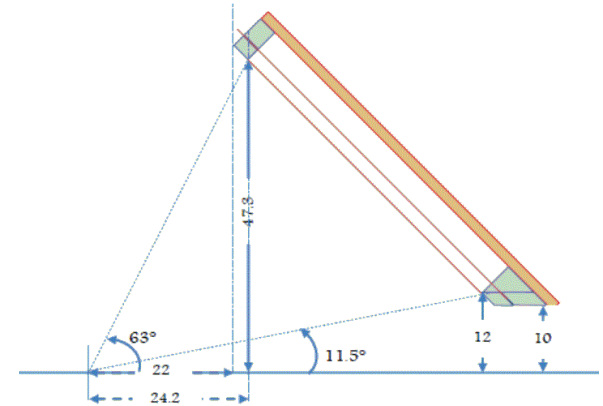
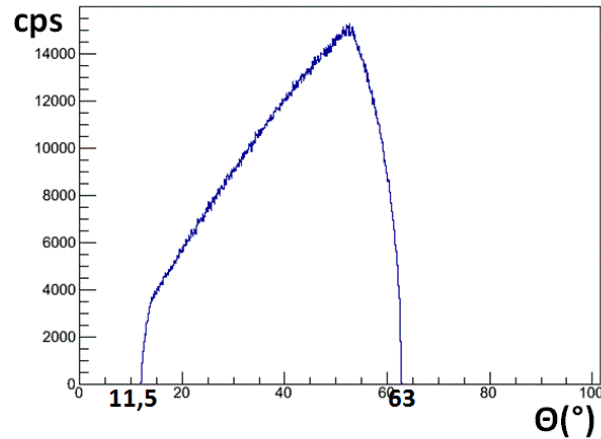
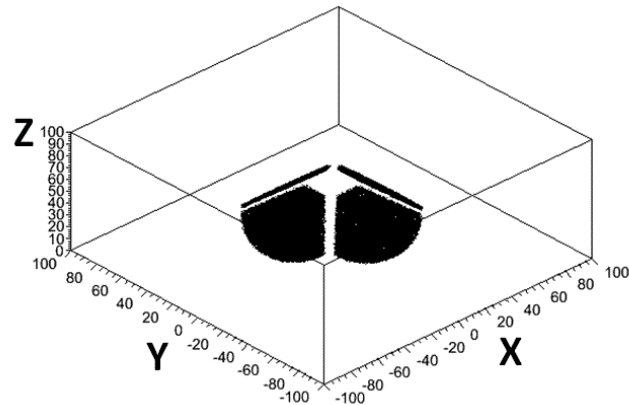
ZOOM

Oslo measurements: results

(neutron-induced) $^{237}\text{Np} + n \rightarrow ^{238}\text{Np}^* \leftarrow ^{238}\text{U}(^3\text{He},t)$ (surrogate)



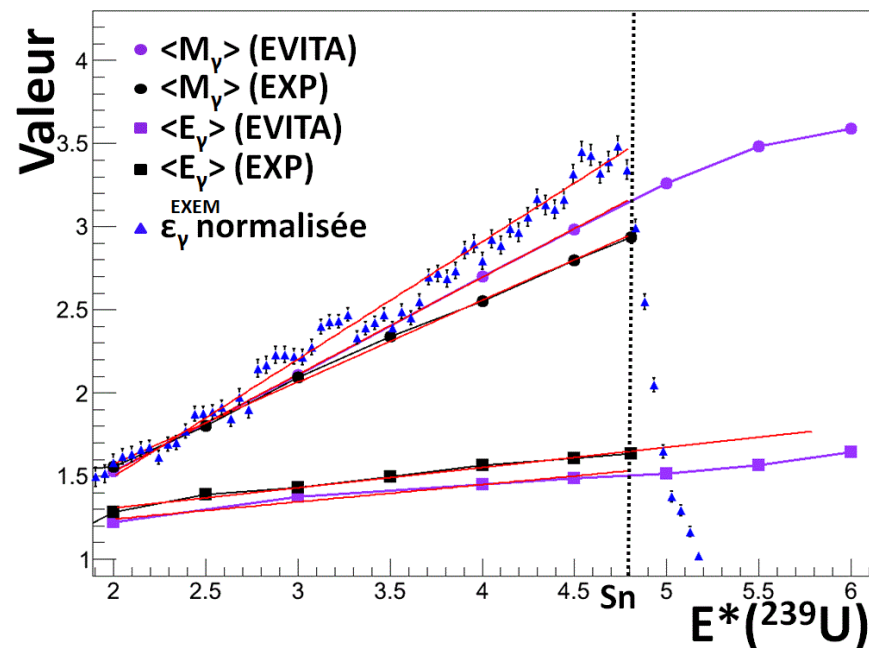
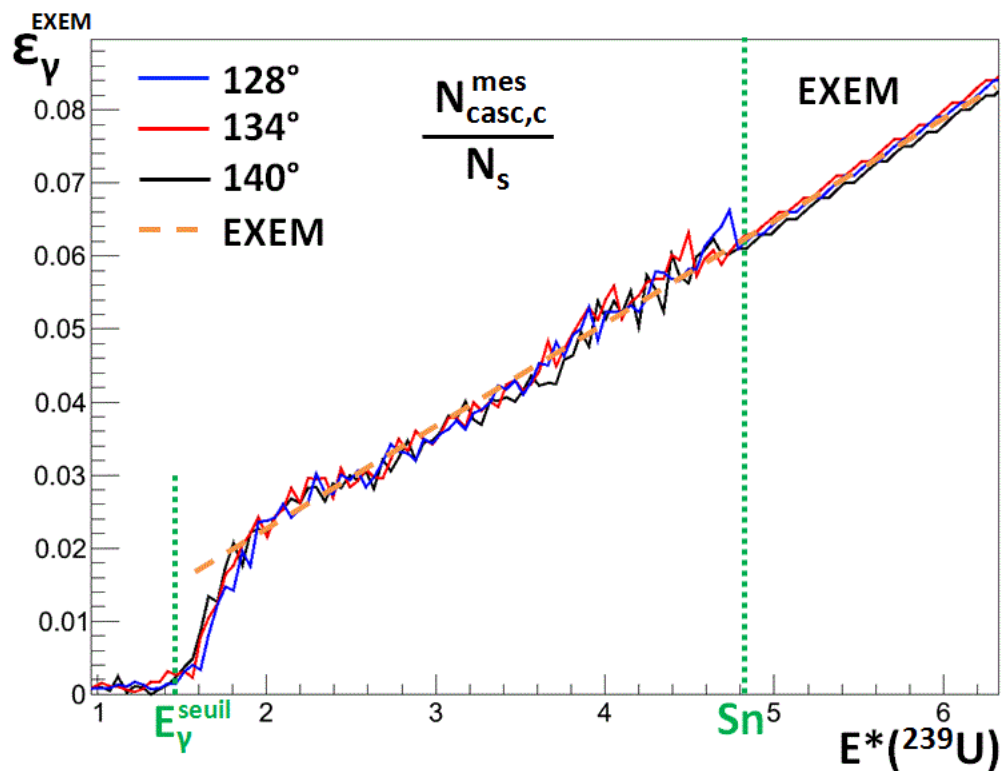
Fission detection efficiency



Gamma detection efficiency

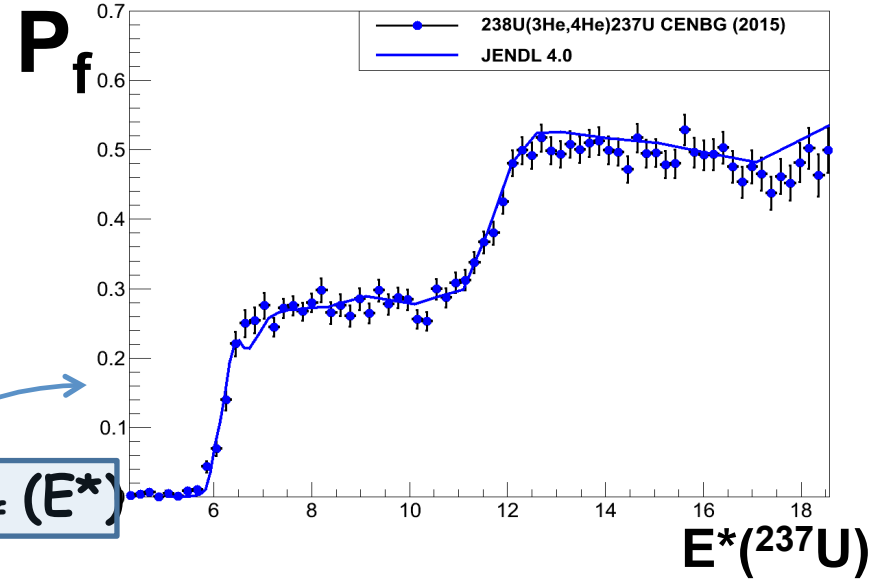
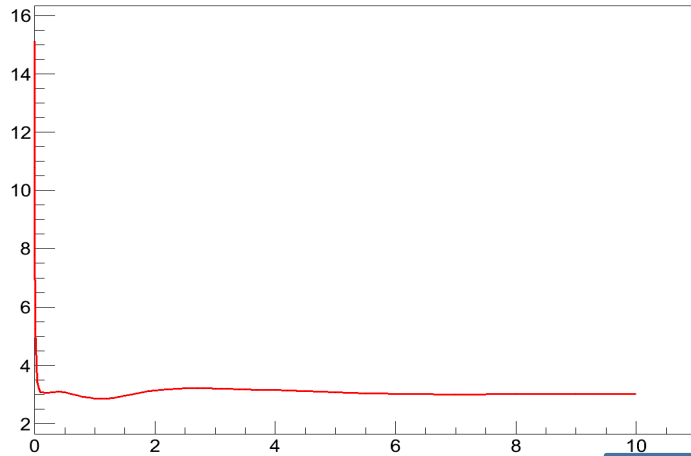
$$P_Y(E^*) = \frac{N_Y(E^*)}{\bar{N}_{\acute{e}j}(E^*) \times \epsilon_Y(E^*)}$$

For $E^* < S_n$, $P_Y = 1 \Rightarrow \epsilon_Y = \frac{N_{\text{ycoinc}}(E^*)}{N_p(E^*)}$

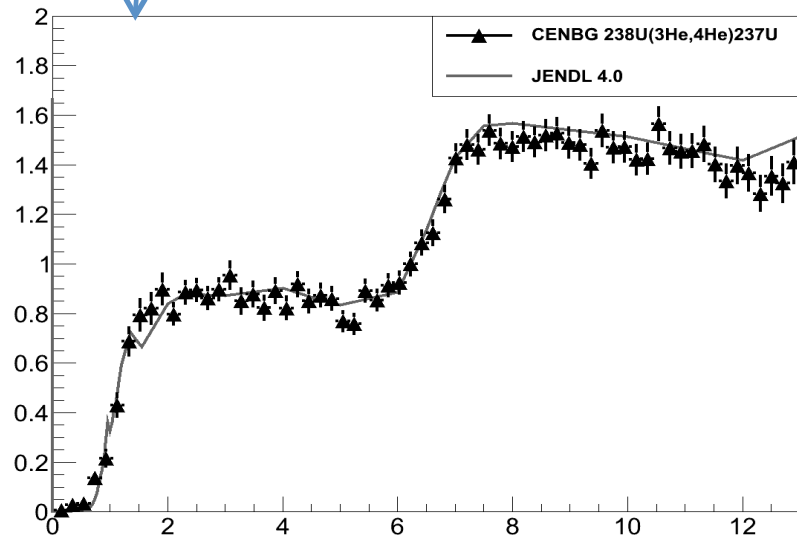


$$\epsilon_Y = f(M_Y, E_Y) \approx k \cdot M_Y \quad (k = \text{constant})$$

Cross section-probability equivalence



$$\sigma_f(E_n) \approx \sigma_{CN}(E_n) \cdot P_f(E^*)$$



Calibration ^3He

