

Exploring gamma strength through Time-dependent Hartree-Fock

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Gamma Strength via TDHF

Overview

- Strength function
- The link to time-dependent Hartree–Fock
- The Sky3D code
- Examples:
 - Some GMR results in spherical nuclei
 - E1 strength / GDR / PDR
 - Pygmy in Se-76

The strength function

The link between the strength function and TDHF

Strength function for a one-body operator given by

$$S(E) = \sum_{\nu} \left| \langle \nu | \hat{M} | 0 \rangle \right|^2 \delta(E - E_{\nu}).$$

Make Fourier representation of delta function

$$S(E) = \frac{1}{2\pi} \int_{-\infty}^{\infty} dt \sum_{\nu} \left| \langle \nu | \hat{M} | 0 \rangle \right|^2 e^{i(E - E_{\nu})t}.$$

Split integral over two parts, with infinitesimal convergence factors for contour integration:

$$S(E) = \frac{1}{2\pi} \sum_{\nu} \left| \langle \nu | \hat{M} | 0 \rangle \right|^2 \left(\int_{-\infty}^0 dt e^{i(E - E_{\nu} - i\delta)t} + \int_0^{\infty} dt e^{i(E - E_{\nu} + i\delta)t} \right)$$

Perform integration to get:

$$S(E) = \frac{-i}{2\pi} \sum_{\nu} \left| \langle \nu | \hat{M} | 0 \rangle \right|^2 \left(\frac{1}{E - E_{\nu} - i\delta} - \frac{1}{E - E_{\nu} + i\delta} \right)$$

The strength function

The link between the strength function and TDHF

Can put the energy denominators inside matrix element:

$$S(E) = -\frac{i}{2\pi} \sum_{\nu} \langle 0 | \hat{M} \left(\frac{1}{E - H - i\delta} - \frac{1}{E - H + i\delta} \right) | \nu \rangle \langle \nu | \hat{M} | 0 \rangle$$

Now can get rid of unit dyadic, leaving no explicit dependence on excited state

$$S(E) = \frac{i}{2\pi} \langle 0 | \hat{M} \left(\frac{1}{H - E + i\delta} - \frac{1}{H - E - i\delta} \right) \hat{M} | 0 \rangle.$$

And use Dirac's formula to write this in a compact spectral representation as

$$S(E) = -\frac{1}{\pi} \text{Im} \langle 0 | \hat{M} \frac{1}{H - E + i\delta} \hat{M} | 0 \rangle$$

Linear response theory gives this matrix element as equal to

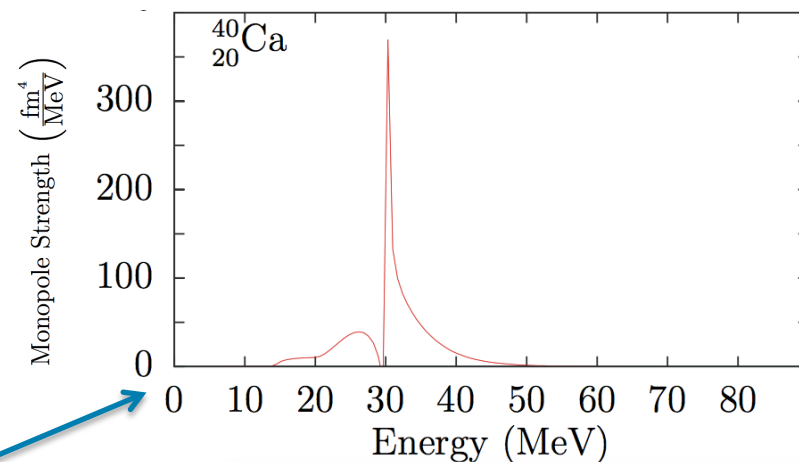
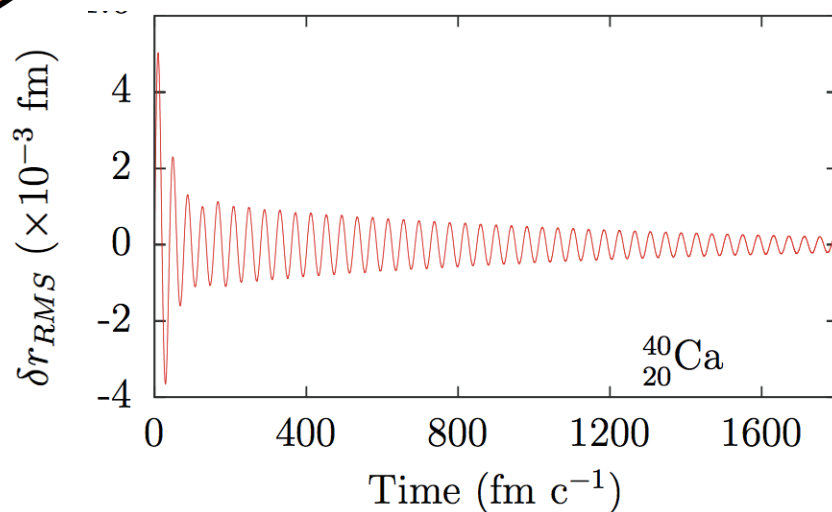
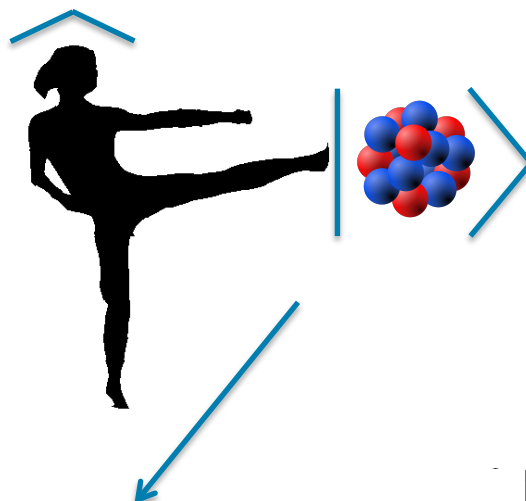
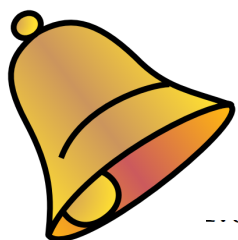
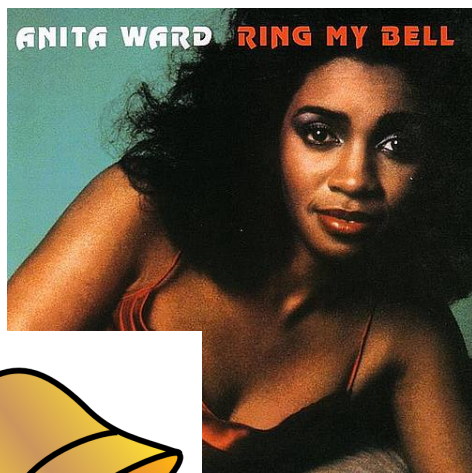
$$\int d^3x m(x) \frac{\delta \langle \rho(x, \omega) \rangle}{f(\omega)}$$

Fourier transform of time-dependent density fluctuations

F-T of time profile of probing kick

In practice for TDHF

(a very schematic guide)



An advertisement

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The TDHF code Sky3D[☆]



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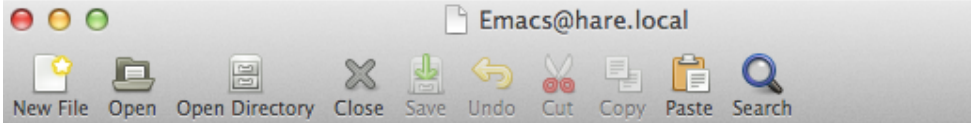
^d Department of Physics and Astronomy, Vanderbilt University, Nashville, TN 37235, USA

- Paper itself is open access
- Detailed subsidiary manual
- Code from CPC program library (or directly from me)

Sky3d code

An advertisement

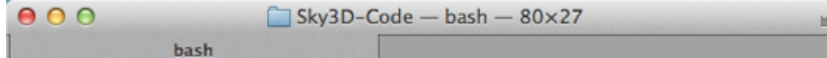
- Modern style Fortran (95/2003)
- Parallelized with OpenMP & MPI
- Very standard conforming –
compiles anywhere (so far)



```

MODULE Forces
  USE Params, ONLY: db,scratch,wflag, hbc, e2, pi
  IMPLICIT NONE
  SAVE
  ! Record defining pairing
  TYPE Pairing
    REAL(db) :: v0prot,v0neut,rho0pr
  END TYPE Pairing
  ! Record defining Skyrme force
  TYPE Force
    CHARACTER(8) :: name
    INTEGER :: ex,zpe
    REAL(db) :: h2m(2)
    REAL(db) :: t0,t1,t2,t3,t4
    REAL(db) :: x0,x1,x2,x3,b4p
    REAL(db) :: power
    TYPE(Pairing) :: vdi ! volume-delta
    TYPE(Pairing) :: ddi ! density-dependent delta
  END TYPE Force
  ! include predefined forces
  INCLUDE 'forces.data'
  ! now the structure used in the run itself
  INTEGER :: ipair
  TYPE(Force) :: f ! force actually used
  TYPE(Pairing) :: p !pairing parameters actually used
  ! charge and mass number in static case for pairing
  REAL(db) :: h2ma ! average h2m over p and n, used in some places
  REAL(db) :: nucleon_mass
  ! derived "b" and Slater coefficients
  REAL(db) :: b0,b0p,b1,b1p,b2,b2p,b3,b3p,b4,b4p,slate
CONTAINS
  SUBROUTINE read_force
    CHARACTER(8) :: name,pairing

```



```

[~/codes/sky3d/Sky3D-Code] hare $ make
gfortran -O3 -c params.f90
gfortran -O3 -c forces.f90
gfortran -O3 -c grids.f90
gfortran -O3 -c fourier.f90
gfortran -O3 -c trivial.f90
gfortran -O3 -c levels.f90
gfortran -O3 -c densities.f90
gfortran -O3 -c sequential.f90
gfortran -O3 -c moment.f90
gfortran -O3 -c coulomb.f90
gfortran -O3 -c meanfield.f90
gfortran -O3 -c inout.f90
gfortran -O3 -c twobody.f90
gfortran -O3 -c fragments.f90
gfortran -O3 -c pairs.f90
gfortran -O3 -c energies.f90
gfortran -O3 -c static.f90
gfortran -O3 -c external.f90
gfortran -O3 -c dynamic.f90
gfortran -O3 -c user.f90
gfortran -O3 -c main3d.f90
gfortran -O3 -o sky3d.seq params.o grids.o levels.o fourier.o forces.o sequen\
al.o inout.o coulomb.o trivial.o densities.o fragments.o twobody.o energies.o st\
atic.o meanfield.o dynamic.o pairs.o moment.o main3d.o user.o external.o -lfftw3\
-llapack
[~/codes/sky3d/Sky3D-Code] hare $

```

(an aside)

PHYSICAL REVIEW E **89**, 033312 (2014)

Extension of the continuum time-dependent Hartree-Fock method to proton states

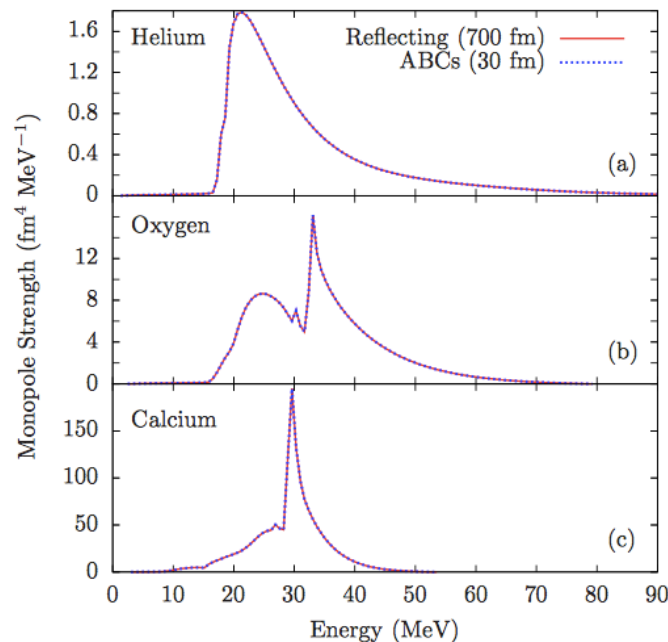
C. I. Pardi* and P. D. Stevenson†

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- Laplace transformation technique
- So far limited to spherical symmetry

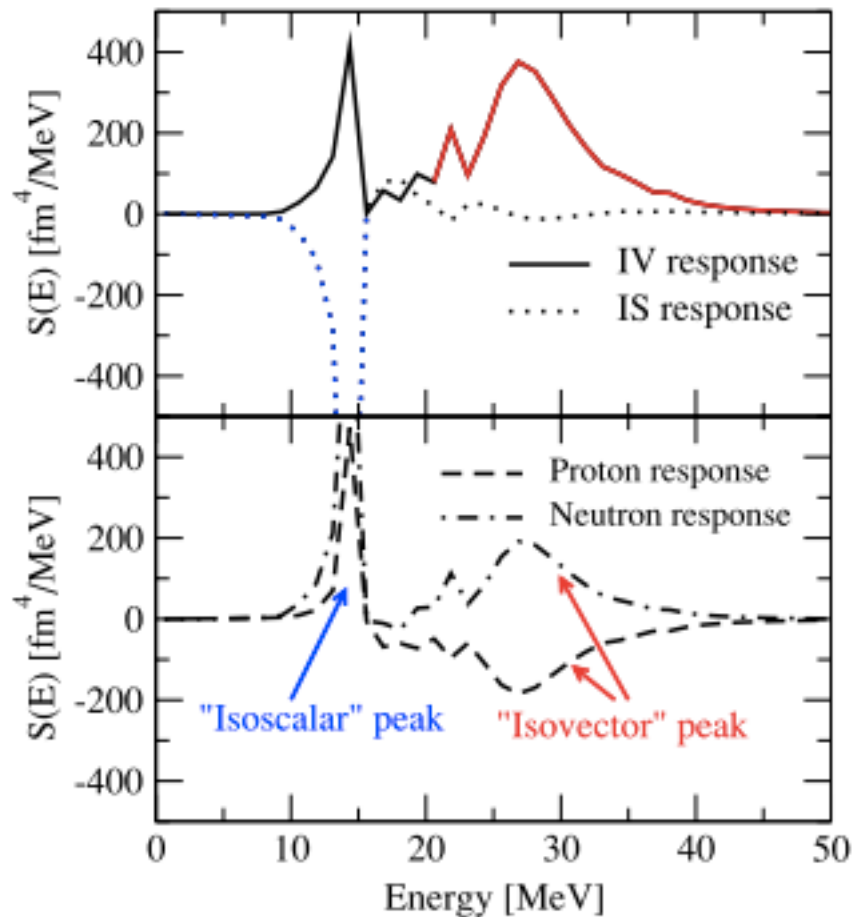
$$Q(R,t) = \int_0^t G_{\kappa,\mu}(R,\tau) \left. \frac{\partial Q(r,t-\tau)}{\partial r} \right|_{r=R} d\tau,$$

where

$$\hat{G}_{\kappa,\mu}(R,s) = \frac{1}{b\sqrt{s}} \left(\frac{W_{\kappa,\mu}(br\sqrt{s})}{\frac{\partial W_{\kappa,\mu}(br\sqrt{s})}{\partial r}} \right) \Big|_{r=R}.$$

A few selected results

Mixed strength functions



$$S(E) = \sum_{\nu} \langle 0 | F | \nu \rangle \langle \nu | \bar{F} | 0 \rangle \delta(E - E_{\nu})$$

F and F-bar different operators – can learn about individual matrix elements

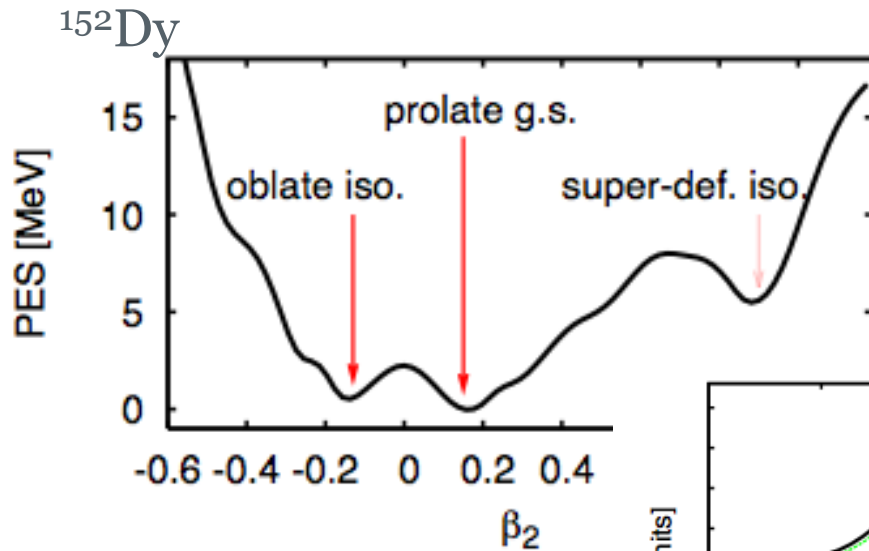
Identify normal modes

Calculations is ²⁰⁸Pb monopole resonances with SkO Skyrme force

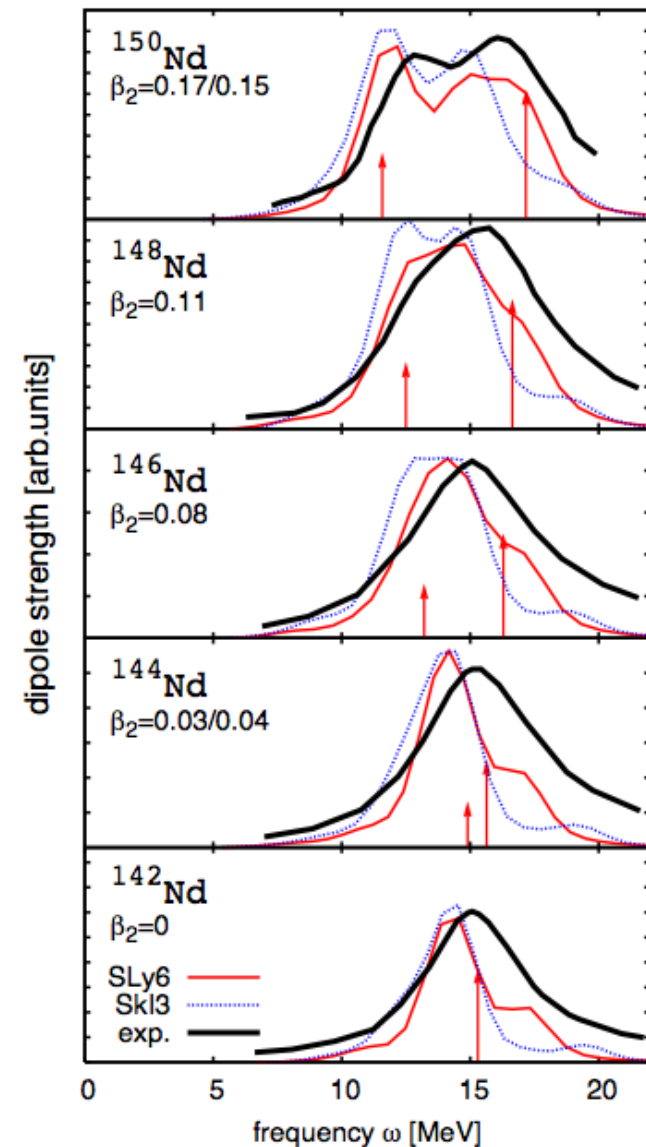
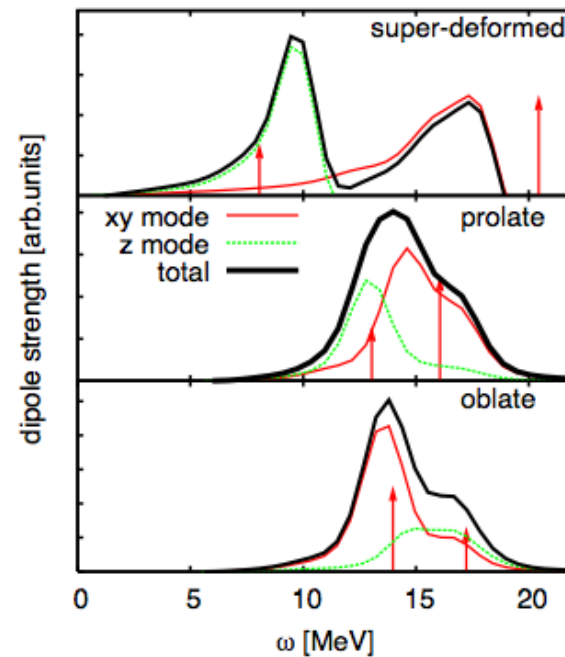
D. Almehed and P. D. Stevenson, J. Phys. G 31, S1819 (2005)

E1 strength

Deformation splitting // GDR on excited states

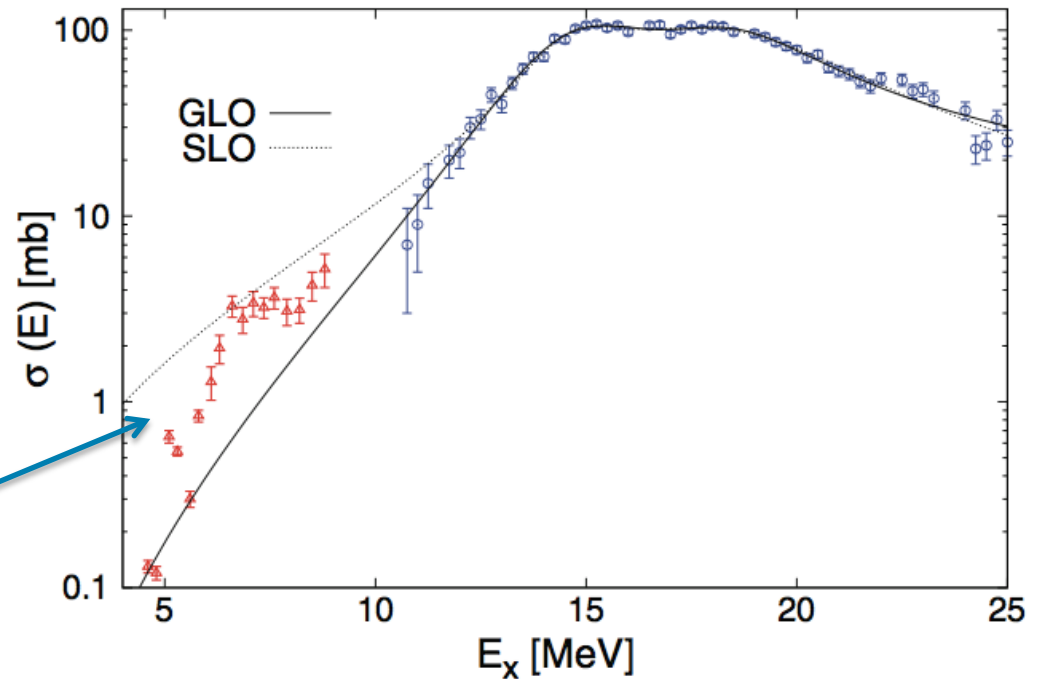


“numerical” width given
by end time of simulation
~2MeV in this case



NRF Technique

^{75}Br	^{76}Br	^{77}Br	^{78}Br	^{79}Br	^{80}Br
^{74}Se	^{75}Se	^{76}Se	^{77}Se	^{78}Se	^{79}Se
^{73}As	^{74}As	^{75}As	^{76}As	^{77}As	^{78}As
^{72}Ge	^{73}Ge	^{74}Ge	^{75}Ge	^{76}Ge	^{77}Ge
^{71}Ga	^{72}Ga	^{73}Ga	^{74}Ga	^{75}Ga	^{76}Ga



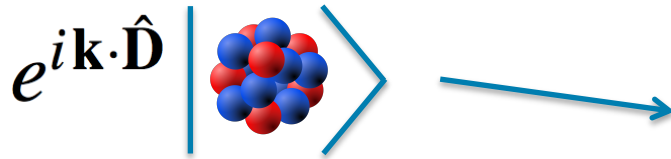
Red triangles: P. M. Goddard, N. Cooper, V. Werner, G. Rusev, P. D. Stevenson, A. Rios, C. Bernards, A. Chakraborty, B. P. Crider, J. Glorius, R. S. Ilieva, J. H. Kelley, E. Kwan, E. E. Peters, N. Pietrella, R. Raut, C. Romig, D. Savran, L. Schnorrenberger, M. K. Smith, K. Sonnabend, A. P. Tonchev, W. Tornow and S. W. Yates, Phys. Rev. C **88**, 064308 (2013)

Above threshold data: P. Carlos, H. Beil, R. Bergre, J. Fagot, A. Leprêtre, A. Veyssièrre and G. Solodukhov, Nucl. Phys. A **258**, 365 (1976)

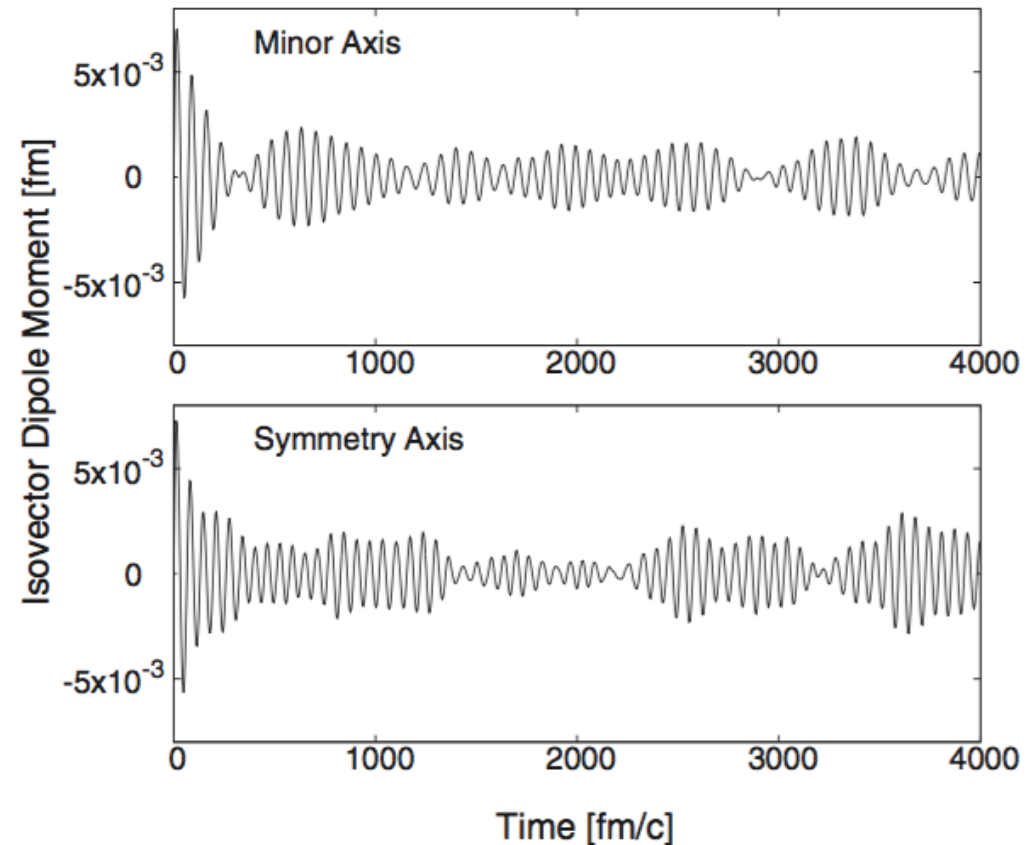
TDHF in ^{76}Se

1. Applying a regular E1 dipole kick

$$\hat{\mathbf{D}} = \frac{N}{A} \sum_{p=1}^Z \mathbf{r}_p - \frac{Z}{A} \sum_{n=1}^N \mathbf{r}_n,$$



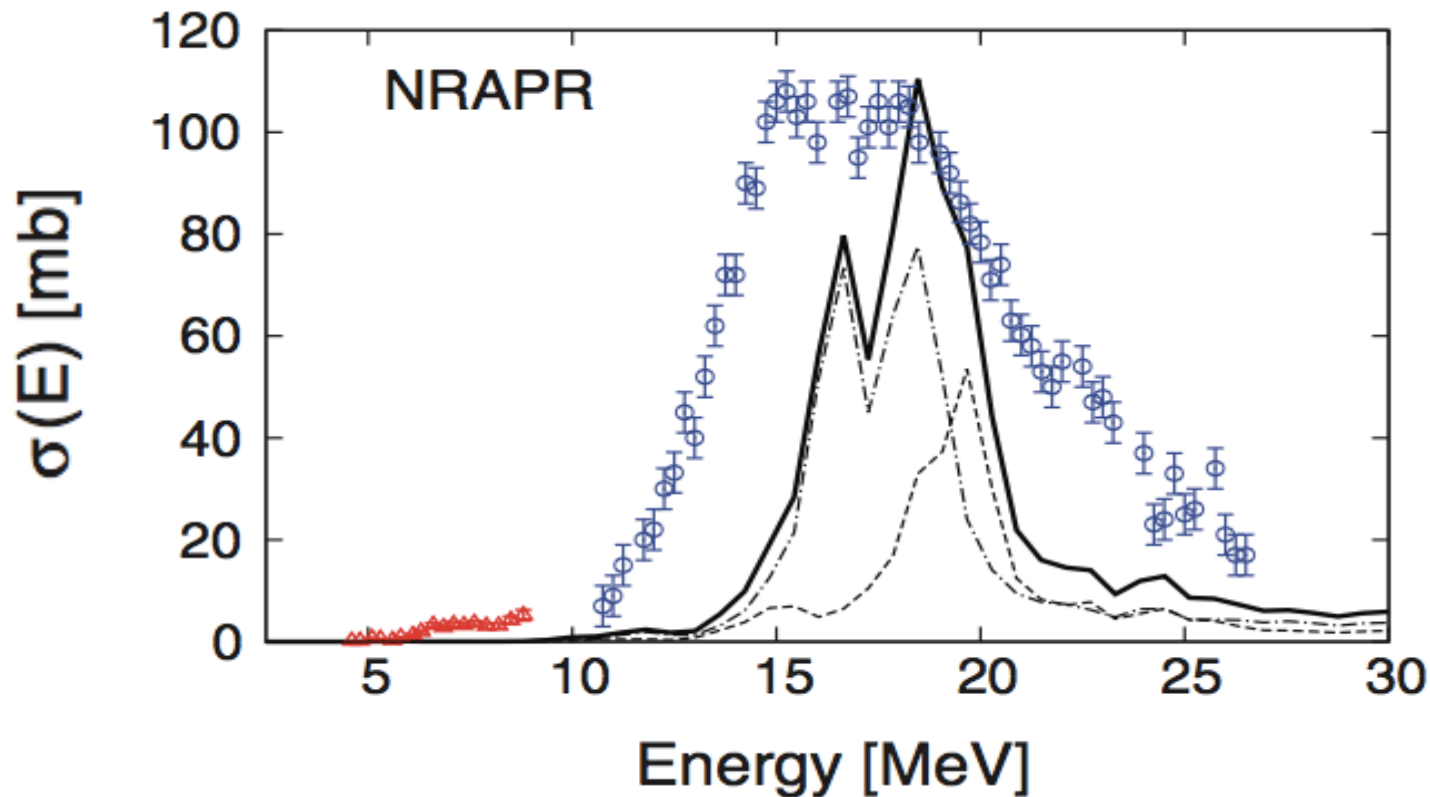
We used Skyrme forces that passed stringent tests for IS and IV nuclear matter properties – M. Dutra, O. Lourenço, J. S. Sá. Martins, A. Delfino, J. R. Stone and P. D. Stevenson, *Phys. Rev. C* **85**, 035201 (2012)



Time-dependent response to dipole kick. Skyrme force NRAPR from A. Steiner, M. Prakash, J. Lattimer and P. Ellis, *Phys. Rep.* **411**, 325 (2005)

Response in ^{76}Se

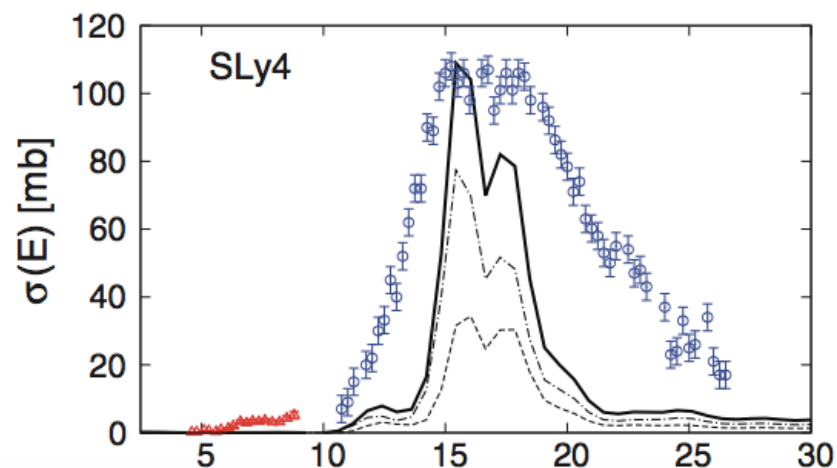
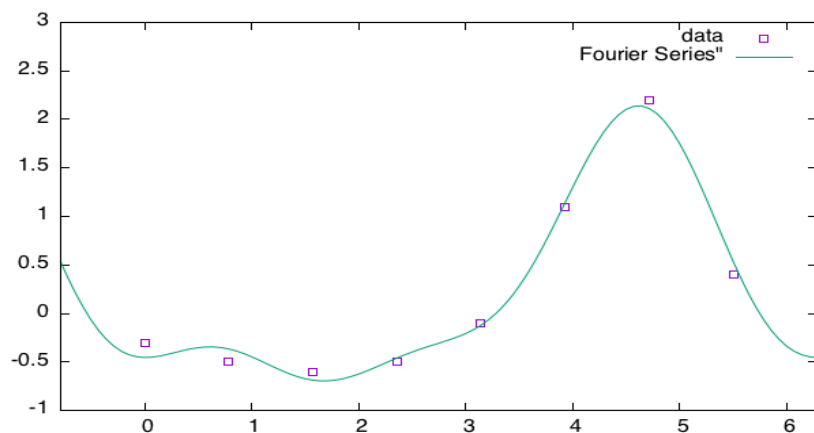
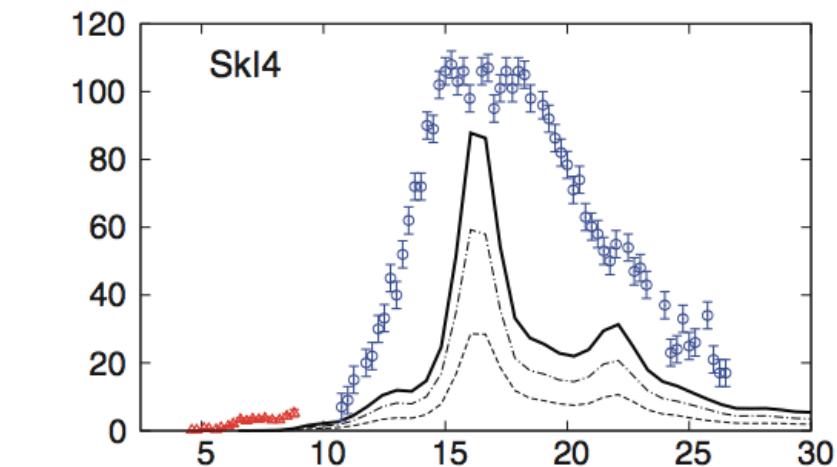
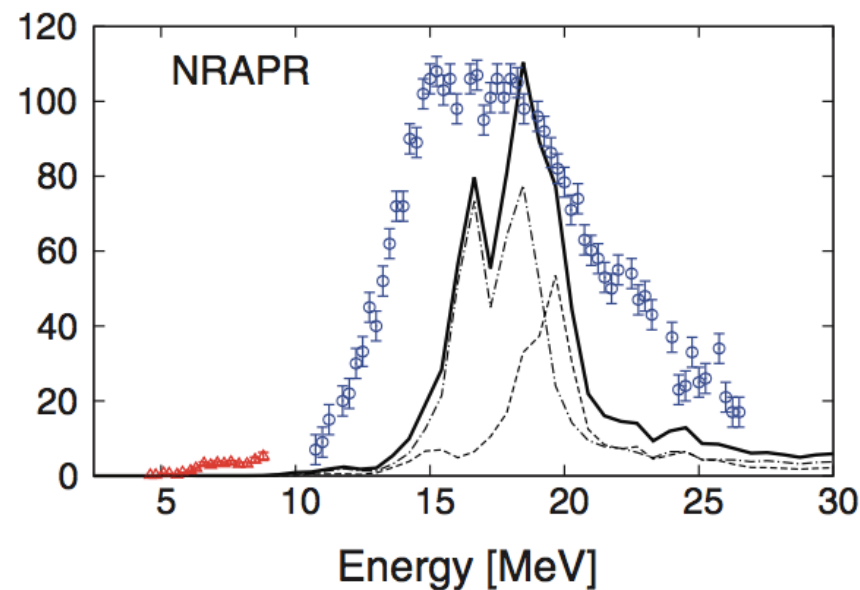
Using NRAPR Skyrme interaction



- Missing spreading width (of course)
- structure from K-splitting and single particle effects

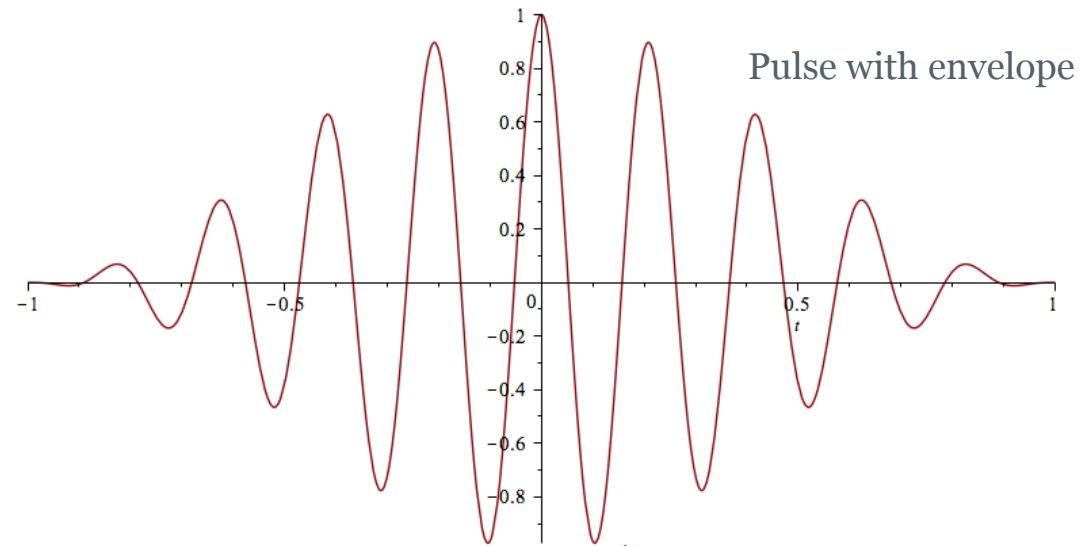
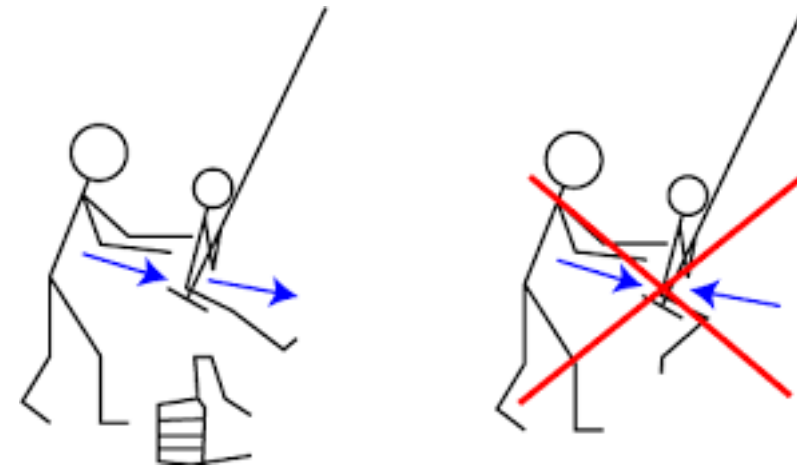
Dependence on Skyrme parameterisation

NRAPR vs SkI4 vs SLy4



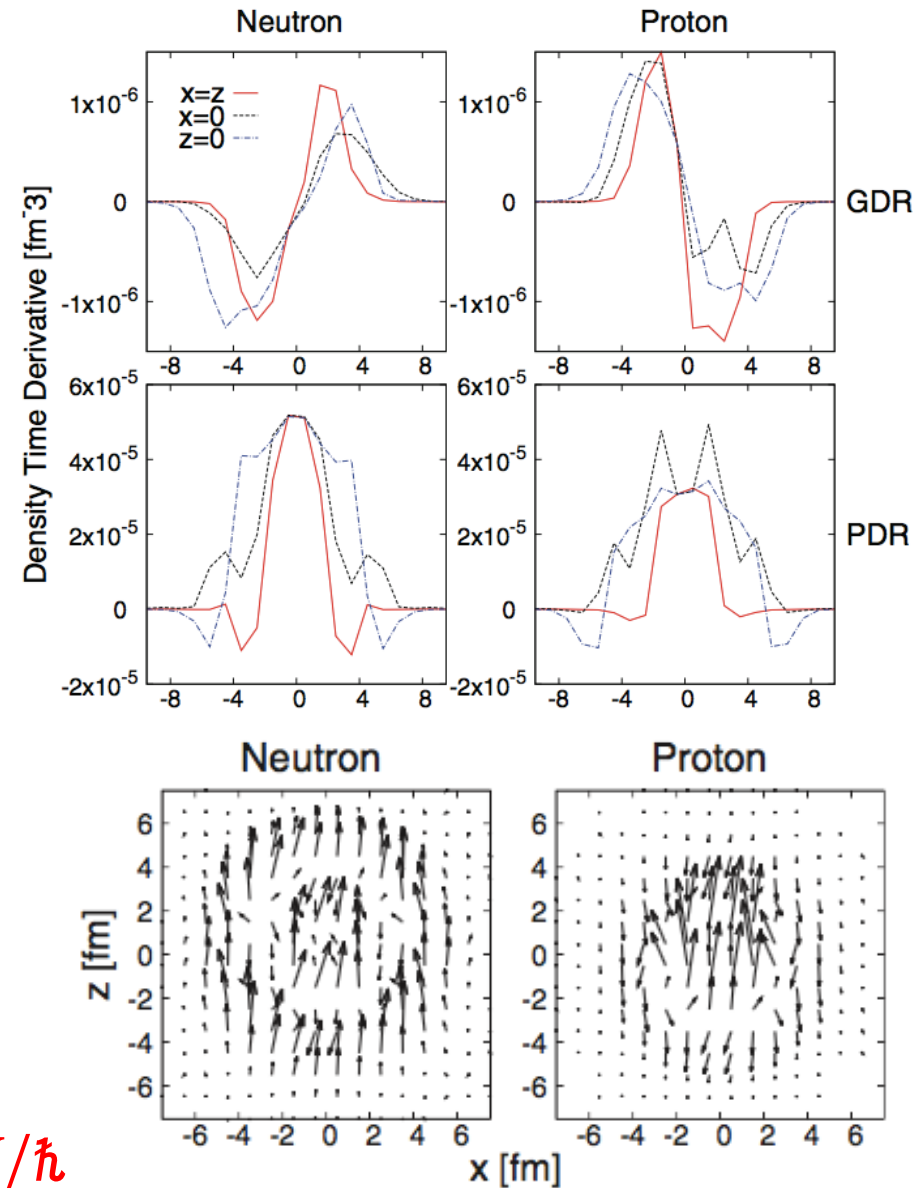
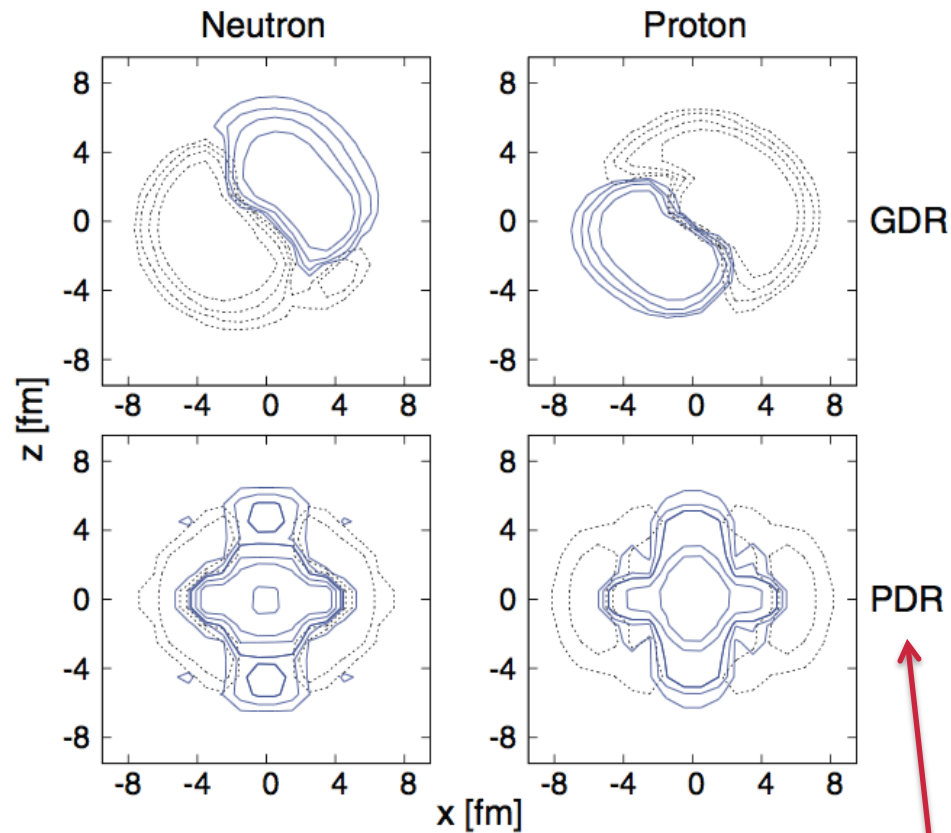
Forcing the nucleus

Driving with a narrowband pulse



$$i\hbar \frac{\partial \rho}{\partial t} = [(\hat{h}_{\text{HF}}[\rho] + f(t)), \rho],$$

Driven density derivatives



$$\omega = 5 \text{ MeV}/\hbar$$

Surface kick

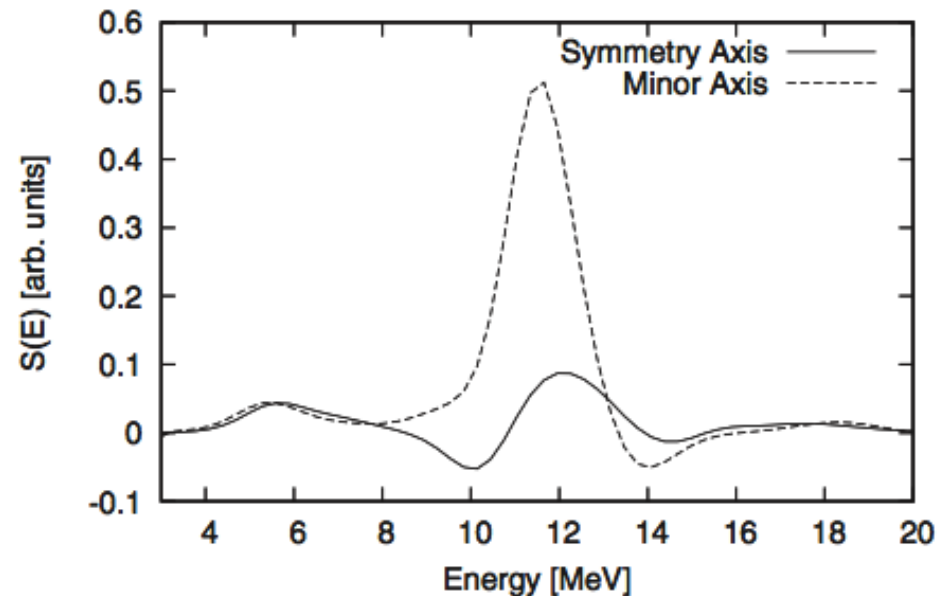
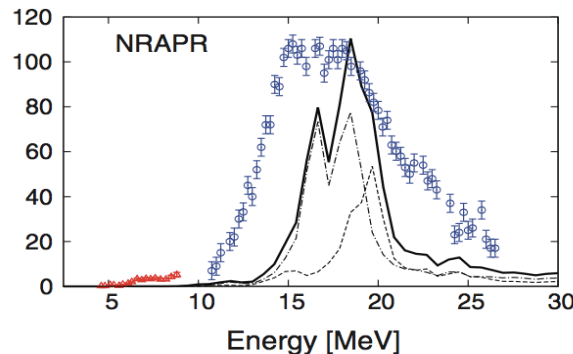
Apply selective dipole boost:

$$\hat{\mathbf{D}}' = \frac{A - \sum v^2}{A} \sum_{i=1}^{\eta} \mathbf{r}_{\text{skin}} - \frac{\sum v^2}{A} \sum_{j=1}^{\zeta} \mathbf{r}_{\text{core}},$$

Sum runs over defined set of core orbitals (the complement of the other sum)

- Strong response around 10–13 MeV
- also spherical bump around 5 MeV
- GDR region unexcited

Sum runs over defined set of skin orbitals – corresponding roughly to 1g_{9/2} neutron orbital (though we have a deformed ground state so correspondence is not exact)



Can initialize TDHF code as collision & study resulting resonance

$^{32}\text{S} + ^{48}\text{Ti}$ @ 66MeV CM



... or look at time-evolution
of fission fragments, or
initialise $t=0$ configuration
as a cluster state, or ...

👏 TDHF is a powerful and
general method. With some
imagination can explore
wide range of nuclear
response (of a one-body
nature)

Thank you:

P. M. Goddard, University of Surrey

J. A. Maruhn, University of Frankfurt



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